



43	1.4.2	Surface area determination
44	1.4.3	Electron microprobe analysis
45	1.4.4	X-Ray diffraction analysis
46	1.4.5	UV-Vis diffuse reflectance spectroscopy
47	1.4.6	Scanning electron microscopy
48	1.5	Chemicals
49		
50	<b>2</b>	<b>Supporting results and discussion</b>
51		
52	2.1	Ozone uptake by Ti-containing minerals under dark conditions
53	2.1.1	Suggested mechanism for ozone uptake by Ti and Ti-bearing minerals
54	2.1.2	Comparison of uptake coefficients to previous literature
55	2.1.3	Influence of relative humidity on ozone uptake
56	2.2	Ozone uptake by TiO <sub>2</sub> under illuminated conditions
57	2.2.1	Mechanistic discussion for photoenhanced ozone uptake by TiO <sub>2</sub>
58	2.2.2	Comparison of uptake coefficients to previous literature
59	2.2.3	Influence of relative humidity on ozone uptake
60	<b>3</b>	<b>References</b>
61		
62	<b>4</b>	<b>Supporting figures</b>
63		
64	S1	Photochemical coated-wall flow tube reactor schematic
65	S2	NO <sub>2</sub> photolysis frequency ( $J_{NO_2}$ ) as a function of number of lamps
66	S3	Representative reaction profiles of ozone with Ti-containing minerals
67	S4	Representative reaction profiles of ozone with TiO <sub>2</sub> minerals
68	S5	Reaction profiles of ozone with titanite 1 and phlogopite as a function of relative
69		humidity
70	S6	Photographs of natural mineral samples prior to grinding
71	S7	Powder diffraction data comparison for anatase samples
72	S8	XRD Sigma purity check for commercial anatase
73	S9–10	XRD diffractogram fits for mineral samples
74	S11	UV-Vis diffuse reflectance spectra
75	S12	Scanning electron microscope images
76	S13	Influence of relative humidity on light ozone BET uptake coefficients ( $\gamma_{BET}$ ) for TiO <sub>2</sub>
77		minerals
78	S14	BET uptake coefficients ( $\gamma_{BET}$ ) under dark and illuminated conditions for ozone by Ti-
79		containing minerals
80		
81	<b>5</b>	<b>Supporting tables</b>
82		
83	S1	Photochemical coated-wall flow tube parameters
84	S2	Composition of TiO <sub>2</sub> /SiO <sub>2</sub> suspensions (for studies of TiO <sub>2</sub> photochemistry)
85	S3	Summary of mineral sample pretreatments
86	S4	Specific surface areas (BET) of samples

87	S5-14	Electron microprobe data for natural minerals
88	S15	XRD table for Pawley fit for minerals
89	S16-20	XRD tables for single crystal diffraction

## 90 **1 Experimental details and sample characterization**

### 91 **1.1 Photochemical coated-wall flow tube reactor**

92 All experiments were conducted in a newly constructed photochemical coated-wall flow  
93 tube reactor system under ambient temperature and pressure conditions. A schematic of  
94 the system is provided in **Figure S1** and the parameters of the flow tube are shown in  
95 **Table S1**.

96 The flow tube itself is constructed of Pyrex 7740 (34.4 cm length, 1.6 cm id, 0.7 cm  
97 thickness water jacket). Insert tubes, also constructed from Pyrex 7740 (20.0 cm length,  
98 1.05 cm id), were coated with the mineral sample of interest and placed inside the flow  
99 tube. Mounted in the center of a black plastic box, the flow tube is surrounded by four  
100 Sylvania 25W BL UV-A lamps (300–410 nm,  $\lambda_{\text{max}}$ : 356 nm). To verify that the light intensity  
101 inside the flow tube was atmospherically relevant, the photolysis frequency of NO<sub>2</sub> ( $J_{\text{NO}_2}$ )  
102 was investigated and found to be  $0.0045 \pm 0.0001 \text{ s}^{-1}$  (see **Section 1.2**), which is in the  
103 range of the photolysis rate constants found in the atmosphere<sup>1,2</sup>.

104 The temperature inside the reactor was controlled by a recirculating chiller (RTE-140,  
105 Neslab), which was set to 296.5 K. To avoid any microbial growth inside the chiller that  
106 could potentially absorb UV radiation in the water jacket layer, the common biocide  
107 chloramine-T was added to the chiller water. The chiller water was frequently monitored  
108 to verify that its absorbance spectrum in the UV-A range matched that of deionized water.

109 Dry, purified zero air for experiments was produced using a commercial zero air generator  
110 (747-30 reactor type A, Aadco Instruments). Ozone was generated in the dry air flow using

111 a commercial ozone generator (97-0066-01, UVP) and monitored at the exit of the flow  
112 tube using a photometric ozone analyzer (T400, Teledyne). Acquisition of the ozone data  
113 was accomplished by connecting the ozone analyzer to an analog-to-digital converter (U6,  
114 LabJack) and collecting the data using a custom-built LabVIEW program at 5 s intervals.

115 The flow tube system contains 4 mass flow controllers (MFCs; MC-2SLPM-D/5M, Alicat), 3  
116 of which were used to control the flows of ozone in dry air, dry air, and wet air prior to the  
117 flow tube. The flow of dry air through the ozone generator was held constant at 100 sccm,  
118 and the relative humidity (RH) in the flow tube was varied by changing the flow ratio  
119 between the dry and wet MFCs. Wet air was generated by passing dry air through a water  
120 bubbler. RH in the system was monitored by 3 in-line humidity/temperature sensors  
121 (SHT75, Sensirion), which were placed directly after the zero air generator in order to  
122 make sure the zero air was dry (RH reading of 0.1%), prior to the flow tube, and at the exit  
123 of the flow tube; RH values prior to and at the exit of the flow tube were always  $\pm 1\%$  of the  
124 values reported in our figures and text. A customized Arduino was employed to convert RH  
125 analog signals into a digital output that was collected using a custom-built LabVIEW  
126 program at 5 s intervals.

127 The total flow upstream of the flow tube was set equal to 400 sccm for all experiments.  
128 Since the ozone analyzer requires a flow rate of 750 sccm, a fourth MFC was placed at the  
129 outlet of the flow tube to add an additional mass flow of 500 sccm, with excess flow  
130 directed to exhaust; all reported ozone concentrations were corrected for this dilution.

131 When experiments were not being conducted, the coated-wall flow tube reactor was  
132 continuously flushed with zero air (400 sccm); in addition, the RH of the zero air flow was

133 adjusted as described above at the end of each day to reflect the desired RH for the  
134 subsequent day's experiments.

## 135 **1.2 Determination of NO<sub>2</sub> photolysis frequency ( $J_{NO_2}$ ) for the flow tube reactor**

136 Irradiance inside the flow tube was quantified using NO<sub>2</sub> actinometry<sup>1,2</sup>, in which the  
137 magnitude of production of NO from NO<sub>2</sub> upon illumination is used to determine the  
138 photolysis frequency for NO<sub>2</sub>,  $J_{NO_2}$ .

139 In these experiments, which were conducted at ambient temperature and pressure, NO<sub>2</sub>  
140 (4.93 ppm ± 5% in N<sub>2</sub>, Certified Standard, Praxair) was diluted with N<sub>2</sub> (5.0 grade, Praxair)  
141 upstream of the flow tube to lead to a final concentration of ~370 ppb. A Na<sub>2</sub>CO<sub>3</sub> denuder  
142 was placed upstream of the flow tube to remove any HONO formed in the cylinder  
143 regulator or wetted parts of the MFCs. The denuder consisted of a sandblasted Pyrex tube  
144 (37.5 cm length, 0.635 cm outer diameter), the interior surface of which was coated with  
145 Na<sub>2</sub>CO<sub>3</sub> solution previously prepared by adding 1 g of glycerol and 1 g of Na<sub>2</sub>CO<sub>3</sub> to 50 mL  
146 of methanol. The coating solution was dripped into the Pyrex tube, which was then dried  
147 with N<sub>2</sub> (5.0 grade, Praxair) prior to use.

148 Upon illumination, the increase in NO and decrease in NO<sub>2</sub> were measured using a  
149 chemiluminescence NO<sub>x</sub> (NO + NO<sub>2</sub>) analyzer (T200U, Teledyne). To prevent the photolysis  
150 of NO<sub>2</sub> in undesired areas and control the illuminated volume inside the flow tube, all but  
151 8.5 cm of the length of the flow tube was covered with aluminum foil; in order to ensure the  
152 light intensity within the reactor was identical to that in our dust photochemistry

153 experiments, these experiments were conducted with a Pyrex insert tube placed inside the  
154 flow tube. From these data,  $J_{NO_2}$  was obtained as follows<sup>2</sup>:

$$155 \quad [NO_2]_{avg} = [NO_2]_{final} + \frac{1}{2}\Delta[NO]$$

$$156 \quad J_{NO_2} = \frac{\Delta[NO]}{[NO_2]_{avg}} \times \frac{F}{V} \times \frac{1}{\phi} \times \frac{1}{T}$$

157 Here,  $[NO_2]_{final}$  is the concentration of  $NO_2$  at the exit of the illuminated flow tube,  $\Delta[NO]$  is  
158 the increase in NO concentration upon illumination,  $F$  is the volumetric flow rate in  $L s^{-1}$ ,  $V$   
159 is the illuminated volume of the reactor in L,  $\phi$  represents the quantum yield of NO  
160 production, and  $T$  represents the transmission factor of the lamps through the reactor  
161 walls.

162 The quantum yield of NO from the photolysis of  $NO_2$  in  $N_2$  depends on several subsequent  
163 reactions induced by the formation of NO. Using the principal reactions described in  
164 Zafonte et al.<sup>2</sup>, we derive a quantum yield of 1.63 for NO for our experimental conditions,  
165 which is in good agreement with the values reported by these authors. Finally, we note that  
166 we set the value of  $T$  as unity, because we are interested in measuring the irradiance within  
167 the flow tube rather than the inherent emission profile of the lamps.

168 Experimental protocol for  $J_{NO_2}$  determination involved first establishing a steady  
169 concentration of  $NO_2$  in the dark ( $\sim 360$  ppb) followed by the following illumination  
170 protocol: 20 min light (1 lamp), 20 min dark, 20 min light (2 lamps), 20 min dark, 20 min  
171 light (3 lamps), 20 min dark, 20 min light (4 lamps), and 20 min dark. This procedure was  
172 conducted in triplicate, with results shown in **Figure S2**. A maximum photolysis frequency

173 of  $0.0045 \pm 0.0001 \text{ s}^{-1}$  was obtained for  $\text{NO}_2$  when all 4 lamps were turned on inside the  
174 reactor, which is similar to the photolysis frequency of  $\text{NO}_2$  in the atmosphere and indicates  
175 an atmospherically relevant irradiance inside the flow tube<sup>1,2</sup>.

### 176 **1.3 Experimental details**

#### 177 **1.3.1 Pyrex insert tube preparation**

178 Pyrex insert tubes were cleaned prior to use by first placing each tube in a 1%  $\text{H}_2\text{SO}_4$  bath  
179 for 1 min followed by a concentrated base bath for 1 h. After the base bath, the tubes were  
180 thoroughly rinsed with deionized water and dried in a gravity oven (100L, Fisherbrand™)  
181 at  $324 \pm 5 \text{ K}$ .

182 The acid bath was prepared by adding 10 mL of concentrated  $\text{H}_2\text{SO}_4$  to 990 mL of deionized  
183 water. The base bath was prepared by adding 1 L of 2-propanol to 166 mL of deionized  
184 water, followed by adding 41.7 g of solid KOH pellets and mixing until fully dissolved.

##### 185 **1.3.1.1 Tube coating procedure for natural minerals**

186 Each coated tube was prepared by adding ~50 mg of Ti-containing mineral to a Pyrex  
187 insert tube, followed by several drops of deionized water to form a slurry. The tube was  
188 then stoppered with rubber stoppers (No. 00), rotated until the slurry covered the entire  
189 inner surface of the tube, and placed on a hot dog roller (RHD800 Retro Series Hot Dog  
190 Roller, Nostalgia Electrics) at maximum temperature to allow the mineral sample to dry in  
191 an even layer. After drying, the first 5 cm of one end of the tube and the first 1 cm of the  
192 other end were wiped clean using moistened laboratory wipes (Kimberly-Clark

193 Professional™). The first 5 cm was wiped to allow the ozone-containing gas flow to achieve  
194 laminar flow prior to its interaction with the mineral surface (see **Table S1**); the last 1 cm  
195 was wiped due to the formation of an uneven coating at the tube ends from the use of the  
196 stoppers. The coated tubes were placed in a gravity oven (100L, Fisherbrand™) at  $324 \pm 5$   
197 K overnight prior to use.

198 The mass of mineral inside each Pyrex insert tube was determined by weighing by  
199 difference on an analytical balance (AB265-S/FACT, Mettler Toledo). Coated tubes were  
200 weighed after each experiment; then, they were rinsed with a tap water/Sparkleen™  
201 mixture, distilled water, methanol, and deionized water, placed in a gravity oven (100L,  
202 Fisherbrand™) at  $324 \pm 5$  K overnight to dry, and weighed again.

### 203 **1.3.1.2 Tube coating procedure for TiO<sub>2</sub> minerals (commercial and natural)**

204 For both the commercial and natural TiO<sub>2</sub> samples, almost complete depletion of ozone was  
205 observed upon illumination, which resulted in a diffusion correction of several orders of  
206 magnitude for  $\gamma$  (as discussed in main text of the manuscript). Mixing of TiO<sub>2</sub> with  
207 photochemically inert SiO<sub>2</sub> has been previously been used to mitigate the effect of  
208 significant depletion of ozone from the gas-phase reservoir upon exposure to illuminated  
209 TiO<sub>2</sub><sup>3</sup>; in this study, a similar strategy was employed.

210 TiO<sub>2</sub>/SiO<sub>2</sub> mixed films were prepared by first sonicating SiO<sub>2</sub> in 30 mL of deionized water  
211 for 1 h (2.8 L Ultrasonic Bath, Fisher Scientific) to reduce the agglomeration of SiO<sub>2</sub> in the  
212 coating procedure, then transferring TiO<sub>2</sub> powder to the SiO<sub>2</sub> suspension. To ensure  
213 quantitative TiO<sub>2</sub> transfer, 1 mL of water was used to rinse the weighing tray that

214 contained the TiO<sub>2</sub>. The resultant mixture was stirred with a PTFE stir bar (1 inch,  
215 Fisherbrand™) to ensure homogeneous mixing of TiO<sub>2</sub> and SiO<sub>2</sub> particles; after 10 min, 1  
216 mL was pipetted into a Pyrex insert tube and the same coating procedure described in  
217 **Section 1.3.1.1** was applied. Because each TiO<sub>2</sub> sample type displayed significantly  
218 different reactivity, different TiO<sub>2</sub> loadings were employed for each TiO<sub>2</sub> sample; the  
219 masses of both TiO<sub>2</sub> and SiO<sub>2</sub> used for each sample type are shown in **Table S2**.

### 220 **1.3.2 Experimental protocol**

221 At the beginning of each experiment, a Pyrex insert tube coated with the mineral substrate  
222 of interest was placed in the coated-wall flow tube reactor, the movable injector was  
223 retracted, and the coating was exposed for 30 min to zero air at the experimental RH. Then,  
224 the injector was pushed in past the mineral coating, preventing exposure of the sample to  
225 the gas flow, and the ozone generator was turned on. After 45 min, at which point a stable  
226 ozone concentration was achieved, the injector was pulled back and samples were exposed  
227 to ozone. After 1 h of dark exposure, samples were illuminated for 1 h; after an additional  
228 30 min of dark exposure, the injector was again pushed in in order to verify that there was  
229 no drift in ozone concentrations throughout the experiment; after 30 min, the ozone  
230 generator was turned off and ozone concentrations were recorded for an additional 5 min.  
231 The average ozone concentrations over the last 5 min in the light, dark, and with the  
232 injector pushed in were used in the calculations described in main text of the manuscript.  
233 Experiments were performed in triplicate for all minerals at 25% RH; selected reaction  
234 profiles for each mineral sample are shown in **Figures S3–4**. For some minerals,

235 experiments were also performed as a function of RH (here, experiments were only  
236 performed once at each RH value); representative reaction profiles are shown in **Figure S5**.

### 237 **1.3.3 Considerations in calculations of uptake coefficients**

238 In this work, we quantify the reactivity of ozone with Ti-containing minerals using a  
239 quantity known as an uptake coefficient ( $\gamma$ )<sup>4</sup>, which is defined as the number of ozone  
240 molecules taken up by the surface divided by the total number of collisions of ozone  
241 molecules with the surface. Previous studies of trace gas–dust interactions have reported  
242 both initial<sup>5-7</sup> and steady-state<sup>5,7-9</sup> uptake coefficients; because the latter is most relevant  
243 for the behavior of dust over extended timescales in the atmosphere, we report it here.

244 In our system, laminar flow conditions (Reynolds number < 2000) and continuum flow  
245 regimes (Knudsen number  $\ll 1$ ) are applicable (see **Table S1**); therefore, at high values of  
246  $\gamma$ , where significant depletion of ozone occurs near the sample surface,  $\gamma_{eff}$  can be limited by  
247 the diffusion of ozone from the centre of the flow tube to the sample surface. In order to  
248 correct for underestimations in  $\gamma_{eff}$  as a result of these radial diffusion limitations, a  
249 correction factor is applied using the CKD method<sup>10</sup> as described in the main text of the  
250 manuscript. We note that we chose mineral sample masses such that reported uptake  
251 coefficients had a maximum correction for radial diffusion of 30% from  $\gamma_{eff}$ .

#### 252 **1.3.3.1 Additional considerations for TiO<sub>2</sub> samples**

253 As described in **Section 1.3.1.2**, in order to minimize the need for excessive corrections for  
254 gas-phase diffusion limitations, we employed TiO<sub>2</sub>/SiO<sub>2</sub> mixed films for all TiO<sub>2</sub> samples.

255 Because SiO<sub>2</sub> itself exhibited non-negligible photochemistry in our experiments,  
256 calculations were slightly more involved for these mixed films.

257 In order to correct for SiO<sub>2</sub> reactivity, we calculated  $\gamma_{BET}$  (25% RH) for 3 SiO<sub>2</sub> samples with  
258 similar masses and used the average  $\gamma_{BET}$  value to back-calculate a  $\gamma_{eff}$  value for the mass of  
259 SiO<sub>2</sub> employed in a given TiO<sub>2</sub>/SiO<sub>2</sub> experiment. From the  $\gamma_{eff}$  thus obtained, we back-  
260 calculated  $k_{obs}$  (s<sup>-1</sup>) for this mass of SiO<sub>2</sub> and subtracted this value from the overall  $k_{obs}$   
261 value for the TiO<sub>2</sub>/SiO<sub>2</sub> mixed film. In this way, we were able to obtain a  $k_{obs}$  value for TiO<sub>2</sub>  
262 alone, which we then used to calculate  $\gamma_{eff}$  and  $\gamma_{BET}$  for TiO<sub>2</sub> as described in the previous  
263 section. Importantly, we used the total surface area of TiO<sub>2</sub> alone in these calculations,  
264 rather than the total surface area of the mixed film. We note that these calculations led to  
265 negligible values for  $k_{obs}$  of TiO<sub>2</sub> alone under dark conditions; as a result, we do not report  
266 dark uptake values for TiO<sub>2</sub>.

267 We note that since radial diffusion limitations are independent of sample type and only  
268 depend on overall loss of ozone onto the surface,  $\gamma_{eff}$  values for the TiO<sub>2</sub>/SiO<sub>2</sub> mixed films  
269 were multiplied by the correction factor ratio associated with  $\gamma_{eff}$  values calculated for the  
270 original mixed films (*i.e.* not for the TiO<sub>2</sub> fraction alone).

## 271 **1.4 Sample treatment, preparation, and characterization**

### 272 **1.4.1 Sample treatment and preparation**

273 Minerals were analyzed for their purity via visual inspection, electron microprobe analysis  
274 (EMPA), and X-ray diffraction (XRD). Impurities were identified for each mineral and  
275 purification treatments were selected accordingly; a summary of all treatments is

276 presented in **Table S3**. Pictures of mineral samples prior to treatment and preparation are  
277 provided in **Figure S6**.

278 Obvious impurities on the outside of mineral samples and impurities that could not be  
279 removed via chemical treatment (*vide infra*) were removed using a Dremel® tool. Rust was  
280 removed by submersing minerals in a 40 g L<sup>-1</sup> solution of SUPER IRON out® (active  
281 ingredients sodium hydrosulfite, sodium metabisulfite, and sodium sulfite) for 30 min.  
282 CaCO<sub>3</sub> was removed by placing minerals in 1 M HCl (phlogopite, 1.5 hours; titanite 1, 1  
283 day). After the removal of impurities, minerals were rinsed with methanol once and with  
284 deionized water 3 times. Minerals that were not subjected to these treatments were also  
285 rinsed in this manner.

286 After minerals underwent purification treatments, they were first broken down using a  
287 steel percussion mortar and pestle. To remove any steel contamination from use of the  
288 percussion mortar, a rare earth magnet wrapped in weighing paper was then passed over  
289 each broken-down sample (this was not possible for ilmenite and hastingsite samples, as  
290 they also displayed magnetic properties). Minerals were then ground using an agate mortar  
291 and pestle until they had a fine, powder-like appearance. To prevent cross contamination,  
292 pure gem-quality Brazilian quartz (from Corinto, Brazil) was used to clean the steel  
293 percussion and agate mortar and pestle between samples by grinding the quartz into a fine  
294 powder; afterward, grinding equipment was rinsed with ethanol.

#### 295 **1.4.2 Surface area determination**

296 Prior to gas adsorption analysis, samples were degassed (activated) on a gas adsorption  
297 sample preparation instrument (Smart VacPrep, Micromeritics); in order to ensure that the  
298 measured surface area is reflective of the surface area presented under our experimental  
299 conditions, degassing was performed in the absence of heat. For activation, each sample  
300 was weighed into a pre-weighed glass sample tube equipped with filler rod and steel frit  
301 ball valve. Samples were subsequently activated by evacuating the sample holder at 5.00  
302 mmHg s<sup>-1</sup> until a pressure level of 1.00 mmHg or lower was achieved; the Smart VacPrep  
303 instrument has a needle on the vacuum port to open the ball valve. Subsequently the  
304 samples were held under unrestricted vacuum for 24 h. The sample holders were then  
305 backfilled with nitrogen gas, removed from the Smart VacPrep instrument, and reweighed.  
306 Gas adsorption isotherm data were collected at 77 K using nitrogen gas as the probe  
307 molecule (3Flex, Micromeritics). Using the MicroActive software suite, the data obtained  
308 were used to determine the Brunauer–Emmett–Teller (BET) accessible surface area by  
309 fitting the linear BET equation and ensuring that the 4-point criterion was met<sup>11</sup>.

### 310 **1.4.3 Electron microprobe analysis**

311 Electron microprobe analysis was conducted on 2 separate dates because minerals were  
312 purchased in 2 groups. The first set of minerals analyzed consisted of anatase 1, rutile,  
313 ilmenite, titanite 1, phlogopite, hastingsite, augite, and epidote. The second set, which was  
314 purchased to enable us to study the reactivity of samples of the same mineral obtained  
315 from different locations, consisted of anatase 2 and titanite 2. All EMPA data can be found  
316 in **Tables S5–14**.

317 **1.4.3.1 Analysis of anatase 1, rutile, ilmenite, titanite 1, phlogopite, hastingsite,**  
318 **augite, and epidote**

319 An electron microprobe (JXA-8900R, JEOL) was used to examine grains of anatase 1, rutile,  
320 ilmenite, titanite 1, phlogopite, hastingsite, augite, and epidote, which had been mounted in  
321 epoxy, polished, and carbon-coated (25 nm thickness) prior to analysis. In addition to back-  
322 scattered electron images, quantitative compositional data were acquired from spot  
323 analyses of the minerals using wavelength-dispersive spectrometry and Probe for EPMA  
324 software<sup>12</sup>. Thirteen elements were measured (Na, Mg, Al, Si, K, Ca, Ti, V, Cr, Mn, Fe, Zn, and  
325 Nb) with the following conditions: 20 kV accelerating voltage, 20 nA probe current, and a  
326 1 µm beam diameter for all minerals except phlogopite, for which a 5 µm beam was used.  
327 Total count times of 20 s were used for both peaks and backgrounds for all elements except  
328 V, Zn, and Nb, for which 30 s was used. The X-ray lines and diffraction crystals were: Na  $K\alpha$ ,  
329 TAP (thallium hydrogen phthalate); Mg  $K\alpha$ , TAP; Al  $K\alpha$ , TAP; Si  $K\alpha$ , TAP; K  $K\alpha$ , PET  
330 (pentaerythritol); Ca  $K\alpha$ , PET; Ti  $K\alpha$ , PET; V  $K\alpha$ , PET; Cr  $K\alpha$ , PET; Mn  $K\alpha$ , LIF (lithium  
331 fluoride); Fe  $K\alpha$ , LIF; Zn  $K\alpha$ , LIF; and Nb  $L\alpha$ , PET. Corrections were applied to V for  
332 interference by Ti, to Cr for interference by V, and to Mn for interference by Cr<sup>13</sup>. X-ray  
333 intensity data were reduced following Armstrong (1995)<sup>14</sup>. The reference standards  
334 consisted of metals, synthetic inorganic materials, and natural minerals<sup>15</sup>.

335 **1.4.3.2 Analysis of anatase 2 and titanite 2**

336 An electron microprobe (SX100, Cameca) was used to examine grains of anatase 2 and  
337 titanite 2, which had been mounted in epoxy, polished, and carbon-coated (25 nm  
338 thickness) prior to analysis. In addition to back-scattered electron images, quantitative

339 compositional data were acquired from spot analyses of the minerals using wavelength-  
340 dispersive spectrometry and Probe for EPMA software<sup>12</sup>. Thirteen elements were  
341 measured (Na, Mg, Al, Si, K, Ca, Ti, V, Cr, Mn, Fe, Zn, and Nb) with the following conditions:  
342 20 kV accelerating voltage, 20 nA probe current, and a 1  $\mu\text{m}$  beam diameter. Total count  
343 times of 30 s were used for both peaks and backgrounds for all elements. The X-ray lines  
344 and diffraction crystals were: Na  $K\alpha$ , TAP; Mg  $K\alpha$ , TAP; Al  $K\alpha$ , TAP; Si  $K\alpha$ , TAP; K  $K\alpha$ , PET; Ca  
345  $K\alpha$ , PET; Ti  $K\alpha$ , PET; V  $K\alpha$ , PET; Cr  $K\alpha$ , PET; Mn  $K\alpha$ , LIF; Fe  $K\alpha$ , LIF; Zn  $K\alpha$ , LIF; and Nb  $L\alpha$ ,  
346 PET. Corrections were applied to V for interference by Ti, to Cr for interference by V, and to  
347 Mn for interference by Cr<sup>13</sup>. However, the correction to V was found to be insufficient,  
348 based on analysis of reference rutile, and so the V data were discarded. The X-ray intensity  
349 data were reduced following Armstrong (1995)<sup>14</sup>. The reference standards consisted of  
350 metals, synthetic inorganic materials, and natural minerals<sup>15</sup>.

#### 351 **1.4.4 X-ray diffraction analysis**

352 Powder X-ray diffraction (XRD) patterns of the ground mineral samples were collected on a  
353 Rigaku Ultima IV diffractometer with a Co  $K\alpha$  radiation source operated at 38 kV and 38  
354 mA. Phase identification was performed using JADE 9.6 software with the 2019 ICDD  
355 Database PDF 4+ and 2018-1 ICSD databases<sup>16,17</sup>. The patterns were analyzed with the  
356 TOPAS academic software package<sup>18</sup>. All phases were analyzed with Pawley refinement  
357 method<sup>19</sup>. X-Ray diffraction data and refinement fits are shown in **Figures S9–10**. The  
358 phases and their statistics from refinement fits are presented in **Table S15**, which also  
359 summarizes the results and lists the cell parameters of the refined phases from the powder  
360 XRD data.

361 The refinements in **Figures S9–10** show good visual fits and excellent refinement statistics  
362 (**Table S15**). Ilmenite had no observed impurity peaks in its diffraction pattern (**Figure**  
363 **S10**). For the rest of the samples, asterisks indicate the most intense impurities observed  
364 with the X-ray diffraction technique. The phlogopite sample exhibited a strong preferred  
365 orientation in the  $0\ 0\ 2n$  direction, which resulted in a strong intensity of the  $0\ 0\ 2n$  peaks  
366 in the powder pattern (**Figure S10**). The samples with major impurity contents  
367 (hastingsite and epidote) were also analyzed with Pawley fit to identify the presence of  
368 secondary phases (see **Figure S10**).

369 Powder diffraction of all anatase ( $\text{TiO}_2$ ) samples revealed minor ( $< 5\%$ ) unidentified  
370 impurities based on a relative intensity comparison (**Figure S7**). The peak at  $\sim 32^\circ\ 2\theta$  could  
371 theoretically be assigned to rutile; the absence of other rutile phase peaks likely reflects the  
372 low abundance of this phase. The commercial anatase sample was purchased from a batch  
373 with a claimed anatase purity of 99.8%; a minor rutile peak was apparent in a quality check  
374 diffraction pattern collected in 2014 (see **Figure S8**). However, **Figure S7** shows multiple  
375 other unidentified peaks in the powder diffraction pattern for this sample. Since most of the  
376 impurities found in all 3 anatase samples gave reflections at the same diffraction angles, a  
377 single crystal was picked from one of the samples (anatase 2) for analysis to check whether  
378 these peaks reflected impurities within the phase itself or the existence of additional  
379 impurity phases. Similarly, titanite 2 (**Figure S10**) features 2 unassigned peaks in its  
380 diffraction pattern that were not present in the diffraction pattern for titanite 1 (**Figure S10**).  
381 In order to check if these peaks originated from the structure of this particular sample of  
382 titanite, a single crystal was picked from the mineral sample and analyzed via single crystal  
383 diffraction.

384 Single crystal diffraction measurements were accomplished using a Bruker PLATFORM  
385 diffractometer equipped with a SMART APEX II CCD area detector and a graphite-  
386 monochromated Mo K $\alpha$  radiation source, using  $\omega$  scans at 8 different  $\phi$  angles with a frame  
387 width of 0.3° and an exposure time of 10 s per frame. Face-indexed absorption corrections  
388 were applied<sup>20</sup>. Structure solution and refinement were carried out using the SHELXTL  
389 (version 6.12) program package<sup>21</sup>. The results of these analyses are displayed in **Tables 16–**  
390 **20**.

391 Atomic coordinates and crystallographic details are comparable to other crystallographic  
392 reports of anatase and titanite crystals; no structural anomalies were observed. Atomic  
393 distances from the crystallographic analysis are similar to multiple single crystal diffraction  
394 reports of these minerals. In the case of anatase 2, a small excess of electron density on Ti  
395 atom (0.4 e<sup>-</sup>/Å<sup>3</sup>) is in a good agreement with the atomic composition (0.92 Ti + 0.08 Fe)  
396 suggested by microprobe analysis (**Table S6**). Additionally, atomic displacement parameters  
397 are slightly higher than what would be expected, suggesting negligible mixing in the metal  
398 site, which is in a good agreement with the microprobe data. The extra peaks (observed on  
399 powder patterns) were absent in the single crystal diffraction data, which suggests that these  
400 peaks indeed belonged to impurity phase(s) and/or K $\beta$  diffraction from the anatase phase  
401 rather than to structural imperfections of the main phase.

#### 402 **1.4.5 UV-Vis diffuse reflectance spectroscopy**

403 In order to determine the extent to which each Ti-containing mineral sample absorbs light  
404 within the transmission range of the UV-A lamps in the photoreactor and to obtain their  
405 band gaps, diffuse reflectance spectra were collected using a UV-Vis-NIR

406 spectrophotometer (Cary, 5000) equipped with a diffuse reflectance accessory.  
407 Spectralon® was used as 100% reflectance standard and spectra were measured from  
408 200–800 nm for all minerals except ilmenite, the spectrum of which was measured to 1200  
409 nm, as its absorbance extended into the NIR region. Diffuse reflectance data were  
410 transformed using the Kubelka–Munk function<sup>22</sup>:

411 
$$\frac{K}{S} = \frac{(1 - R)^2}{2R}$$

412 Here,  $K$  is the absorption coefficient,  $S$  is the scattering coefficient, and  $R$  is the reflectance.  
413 Band gaps for Ti-containing minerals were determined by plotting the Kubelka–Munk  
414 function ( $K/S$ ) against eV, where eV corresponds to the kinetic energy of a single electron at  
415 a given wavelength, and then taking the intersection between a linear fit of the reflectance  
416 dip and a horizontal line along the baseline of  $K/S$ <sup>23</sup>. The Kubelka–Munk plots for a  
417 selection of minerals (with associated band gaps, when possible) are presented in **Figure**  
418 **4**; the remainder of the minerals are shown in **Figure S11**.

#### 419 **1.4.6 Scanning electron microscopy**

420 The size and morphology of a select set of mineral samples—commercial anatase, anatase  
421 1, anatase 2, rutile, titanite 1, and titanite 2—were assessed using a field emission scanning  
422 electron microscope (FESEM, Zeiss Sigma). In order to prepare samples for SEM analysis, a  
423 small amount (several mg) of each mineral was sprinkled onto carbon tape, which was then  
424 mounted on an aluminum stub. To reduce the influence of surface charging from the  
425 electron beam, samples were carbon coated using a sputter coater (EM SCD005, Leica).

426 Secondary electron images were collected using in-lens mode at 10 kV at 1000×  
427 magnification, except in the case of commercial anatase, which was obtained at 10,000×  
428 magnification due to its significantly smaller particle sizes compared to the natural ground  
429 minerals. All secondary electron images are presented in **Figure S12**.

## 430 **1.5 Chemicals**

431 Natural minerals were sourced from [www.minfind.com](http://www.minfind.com) and through contact with  
432 individual vendors. Deionized water (18 MΩ) was obtained from a Millipore Synergy UV  
433 ultrapure water system. TiO<sub>2</sub> (anatase, 99.8% trace metal basis), KOH (ACS reagent ≥ 85%,  
434 pellets), and chloramine T trihydrate (Reag. Ph. Eur. grade) were obtained from Sigma  
435 Aldrich. SiO<sub>2</sub> (amorphous fumed, reported surface area 85–115 m<sup>2</sup> g<sup>-1</sup>) was obtained from  
436 Alfa Aesar. Methanol (Optima grade, 99.9%), 2-propanol (certified ACS Plus), Na<sub>2</sub>CO<sub>3</sub>  
437 (anhydrous, certified ACS powder), and glycerol (certified ACS) were obtained from Fisher  
438 Chemicals. Concentrated H<sub>2</sub>SO<sub>4</sub> (reagent grade) and concentrated HCl (reagent grade) were  
439 obtained from Caledon Laboratory Chemicals. SUPER IRON out<sup>®</sup> (Summit Brands) was  
440 purchased from Canadian Tire<sup>®</sup>. Sparkleen<sup>™</sup> was purchased from Fisherbrand<sup>™</sup>. Ethanol  
441 (95%) was obtained from Commercial Alcohols. Except for the natural mineral samples,  
442 which were treated as described in **Section 1.4.1**, all chemicals were used as received.

## 443 **2 Supporting results and discussion**

### 444 **2.1 Ozone uptake by Ti-containing minerals under dark conditions**

445 Although the primary focus of this study was the photochemistry of Ti-containing minerals,  
446 we also investigated ozone uptake by these minerals under dark conditions. As illustrated  
447 in **Figures S3–4**, ozone exhibited time-dependent uptake to most Ti-containing mineral  
448 samples under dark conditions. In particular, for most samples, the magnitude of ozone  
449 loss decreased with increasing sample exposure and eventually reached a plateau value; in  
450 other words, the concentration of ozone at the exit of the flow tube increased with  
451 increasing sample exposure and eventually reached a steady state. We note that in some  
452 cases, steady-state conditions were not reached (*e.g.* phlogopite, titanite 1, anatase 1, and  
453 anatase 2; **Figures S3–4**), as we wanted to keep the total reaction time similar between  
454 experiments; for these minerals, therefore,  $\gamma_{\text{BET}}$  values represent an upper limit.

455 As shown in **Figure S14**,  $\gamma_{\text{BET}}$  values for our samples under dark conditions were  
456 significantly lower than values obtained under illumination. Under dark conditions,  $\gamma_{\text{BET}}$   
457 values spanned almost an order of magnitude, with the highest values calculated for  
458 ilmenite ( $7.6 \times 10^{-7}$ ) and the lowest for epidote ( $(7.5 \pm 1.8) \times 10^{-8}$ ). We note that ilmenite  
459 displayed 2 distinct types of reaction profiles with associated uptake coefficients ( $(6.9 \pm$   
460  $1.3) \times 10^{-7}$ ,  $n = 3$ ;  $(2.4 \pm 0.6) \times 10^{-7}$ ,  $n = 3$ ); although we are not currently able to definitively  
461 explain this behavior, we tentatively suggest that it may reflect variable contributions from  
462 additional minor mineral phases within the sample. If present, these additional mineral  
463 phases may have gone undetected during our sample characterization procedures, as  
464 EMPA results only reflect the elemental abundances in spot-selected areas of mineral  
465 samples and XRD is not always able to detect minor phases in multi-phase samples<sup>24</sup>. To  
466 compare the reactivity of our minerals to natural dust samples, we also measured ozone

467 uptake by Cape Verde dust and Gobi dust; for these samples,  $\gamma_{\text{BET}}$  values of  $(2.2 \pm 0.2) \times 10^{-8}$   
468 and  $(3.6 \pm 0.6) \times 10^{-8}$  were obtained, respectively.

469 Ozone uptake by  $\text{TiO}_2/\text{SiO}_2$  mixed films under dark conditions was indistinguishable from  
470 that by  $\text{SiO}_2$  films, which was itself small. These results agree with the coated-wall flow  
471 tube study of Nicolas et al., which reported minimal uptake of ozone by  $\text{TiO}_2/\text{SiO}_2$  films  
472 under dark conditions<sup>3</sup>. As noted in **Section 1.3.3.1**, because of the substantial  
473 photoreactivity of  $\text{TiO}_2$  samples, our films were prepared using low Ti mass fractions (see  
474 **Table S2**), which implies that  $\text{TiO}_2$  minerals themselves may be measurably reactive  
475 toward ozone under dark conditions; however, this was not the focus of the present study  
476 and is not explored/discussed here.

#### 477 **2.1.1 Suggested mechanism for ozone uptake by Ti and Ti-bearing minerals**

478 The following mechanism has been proposed for the interaction of ozone with metal oxide  
479 surfaces, where SS represents an active surface site (we note that the importance of **R4** is  
480 debated)<sup>25,26</sup>:



485 The saturation of active surface sites via **R1** and **R2** implies that ozone uptake by these  
486 samples will eventually become negligible; however, the mechanism also allows for

487 regeneration of these sites via **R3** and **R4**, which would lead to sustained (*i.e.* catalytic)  
488 ozone loss. Our steady-state ozone concentrations do not return back to their initial values  
489 prior to sample exposure (see **Figures S3–4**), which implies that destruction of ozone is  
490 catalytic in nature to some degree for all of our samples. This catalytic behavior has been  
491 previously observed for several natural mineral dusts<sup>5,9,27</sup>, clays<sup>5,8,29</sup>, and metal  
492 oxides<sup>3,5,28,30,31</sup>, the latter two of which are commonly used as mineral dust proxies.

493 The surface sites (SS) presented in **R1–4** represent active Lewis acid sites (*i.e.* surface  
494 defects in the form of oxygen vacancies) that can promote ozone dissociation and resultant  
495 formation of surface-bound oxygen atoms; in fact, in the case of volcanic ash, ozone has  
496 been used as a molecular probe for Lewis acid sites<sup>32</sup>. The adsorption of ozone at these  
497 Lewis acid sites has been studied using many surface spectroscopic techniques, including  
498 FT-IR<sup>33</sup>, XPS<sup>34</sup>, Raman<sup>26,35,36</sup>, and DRIFTS<sup>37</sup>. The degree to which the catalytic destruction of  
499 ozone occurs depends greatly on the strength of these Lewis acid sites. For example,  
500 previous studies have found that the catalytic uptake of ozone by Fe<sub>2</sub>O<sub>3</sub> is greater than by  
501 Al<sub>2</sub>O<sub>3</sub>, which is in turn greater than by SiO<sub>2</sub><sup>5</sup>. This trend has been attributed to Fe<sub>2</sub>O<sub>3</sub> having  
502 the weakest Lewis acid sites—and, by extension, the weakest SS–O<sub>2</sub> bonds—which allows  
503 for greater regeneration of SS<sup>30</sup>.

504 We suggest that similar considerations may be relevant for our sample set, which includes  
505 minerals with different chemical compositions and crystal structures (*i.e.* that observed  
506 reactivity differences may arise from differences in strength of Lewis acid sites). For  
507 example, ilmenite’s elevated dark reactivity compared to other samples may reflect its  
508 elevated Fe content (~45%, **Table S8**) and consequent availability of Fe cations at its

509 surface. In addition, titanite 1's reactivity ( $\gamma_{\text{BET}} = (1.29 \pm 0.02) \times 10^{-7}$ ) compared to titanite 2  
510 (uptake negligible compared to controls, and therefore not reported) may reflect the 4×  
511 larger amount of Fe detected in the former sample (**Tables S9–10**).

512 Fe can easily substitute for Ti in mineral structures due to its similar cation size and  
513 charge<sup>38</sup>. Depending on formation processes and conditions (*e.g.* crystallization  
514 environment), therefore, we expect titanium-containing minerals in the environment to  
515 have different abundances of these elements in their crystal structures, and ultimately  
516 exhibit different reactivities; in other words, we anticipate that the results obtained here  
517 are not specific for titanite but rather reflective of broader variability in environmental  
518 samples.

### 519 **2.1.2 Comparison of uptake coefficients to previous literature**

520 Previous laboratory studies investigating the uptake of ozone by natural mineral dusts and  
521 mineral dust proxies have been conducted using coated-wall flow tubes<sup>8,25,29,39</sup>,  
522 environmental chambers<sup>30,31</sup>, and Knudsen cells<sup>5,27,28</sup>. Detailed descriptions of these  
523 systems are provided in a review by Kolb et al.<sup>40</sup>, but the most significant difference  
524 between them is that whereas Knudsen cells are operated under molecular flow conditions  
525 (*i.e.* under vacuum), coated-wall flow tubes and environmental chambers can be operated  
526 at atmospheric pressures.

527 Uptake coefficients reported in Knudsen cell studies of dust–ozone interactions are  
528 significantly larger than those obtained in this work. For example, Hanisch and Crowley  
529 reported a steady-state BET uptake coefficient of  $7 \times 10^{-6}$  for the uptake of ozone by

530 Saharan Desert dust<sup>41</sup>, which is ~9–90× larger than the values obtained in our experiments  
531 for our natural Ti-containing mineral samples and 200–300× larger than our Gobi and  
532 Saharan dust samples, respectively. In the following paragraph, we suggest three  
533 possibilities to explain this discrepancy.

534 First, as recently discussed by Lasne et al., the larger ozone uptake coefficients obtained in  
535 Knudsen cell experiments could potentially be explained by more favorable regeneration of  
536 surface sites via decomposition of SS–O<sub>2</sub> (**R3**) at lower operating pressures<sup>8</sup>. Second, as  
537 noted by Coates Fuentes et al., who reported smaller ozone uptake coefficients for clays in  
538 their coated-wall flow tube experiments as compared to previous values from Knudsen cell  
539 experiments, the *in situ* BET surface area of samples may differ from that measured when  
540 samples are present in powder form<sup>29</sup>. A reduction in the actual BET surface area of our  
541 coated tubes with respect to the values determined for dry samples would also lead to an  
542 underestimation in our calculated  $\gamma_{\text{BET}}$  values. In the case of ozone uptake by hydrocarbon  
543 soot, *in situ* gas adsorption analysis<sup>42</sup> has previously been used to address these issues;  
544 however, this approach would be challenging for our samples, which have significantly  
545 lower specific surface areas. Third, we show in **Section 2.1.3** for a subset of our mineral  
546 samples that the uptake of ozone decreases with increasing RH. By extension, we might  
547 expect uptake coefficients to be higher in low-pressure reactors, where adsorbed water  
548 may be more thoroughly removed from samples and the number of available surface sites  
549 for ozone adsorption may be correspondingly increased.

550 Uptake coefficients derived from studies conducted at atmospheric pressure are in better  
551 agreement with our values<sup>3,8,25,29–31</sup>. For example, Mogili et al. investigated ozone uptake

552 onto  $\alpha$ -Al<sub>2</sub>O<sub>3</sub> and  $\alpha$ -Fe<sub>2</sub>O<sub>3</sub> as a function of RH and ozone partial pressure in an atmospheric  
553 chamber and obtained  $\gamma_{\text{BET}}$  values ranging from  $1.9 \times 10^{-9}$  to  $1.3 \times 10^{-7}$  for these samples<sup>30</sup>.  
554 In addition, Coates Fuentes et al. studied ozone uptake onto kaolinite, a clay used as a  
555 mineral dust proxy, using a coated-wall flow tube and found a  $\gamma_{\text{BET}}$  value of  $1.9 \times 10^{-8}$  at RH  
556 <1%<sup>29</sup>; similar results were obtained by Lasne et al.<sup>8</sup>. Finally, our results for Gobi dust are  
557 similar to those reported in a previous study of the catalytic uptake of ozone by this  
558 substrate in a U-shape fixed-bed reactor<sup>9</sup>.

### 559 **2.1.3 Influence of relative humidity on ozone uptake**

560 Although mineral dust is emitted primarily in arid regions, it can be exposed to  
561 environments of differing RH during atmospheric transport. For example, mineral dust is  
562 primarily transported across the Atlantic Ocean in the Saharan Air Layer, a dry, warm, air  
563 mass; during transport, however, tropical disturbances can lead to vertical mixing of this  
564 layer with underlying cool, moist, tropical Atlantic air and ultimately expose dust to more  
565 humid environmental conditions<sup>43,44</sup>. In addition, the mixing between dry, dust-laden air  
566 flows and more humid air can also occur during dust storm episodes that occur at lower  
567 altitudes<sup>45</sup>. In this context, it is important to understand the influence of water vapour on  
568 the reactivity of our samples toward ozone.

569 Here, we investigated the RH dependence of ozone uptake for all TiO<sub>2</sub> minerals studied  
570 (commercial anatase, anatase 1, anatase 2, and rutile) and for 4 Ti-containing minerals  
571 (titanite 1, titanite 2, phlogopite, and ilmenite). Because we observed minimal loss of ozone  
572 for all TiO<sub>2</sub>/SiO<sub>2</sub> samples under dark conditions, discussion of the influence of RH on ozone  
573 uptake by these samples is presented in **Section 2.2.3** (illuminated conditions). In the case

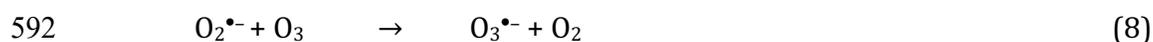
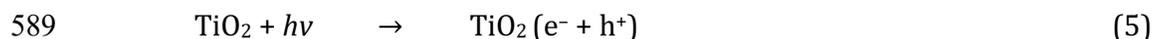
574 of titanite 2, dark reactivity was small at all RH values (results not shown), so trends in  
575 uptake were difficult to discern. In the case of ilmenite, the bimodal reactivity described in  
576 **Section 2.1** precluded reliable interpretation of RH trends. For these reasons, we limit our  
577 discussion to the influence of RH on ozone uptake by titanite 1 and phlogopite.

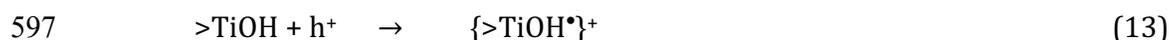
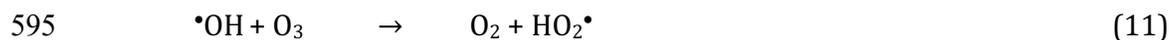
578 As illustrated in **Figure S5**, the loss of ozone at the surface of both titanite 1 and phlogopite  
579 increased with decreasing RH. At lower RH values, the correction factor for radial diffusion  
580 of ozone inside the coated-wall flow tube becomes very large; for this reason, we did not  
581 calculate  $\gamma_{\text{BET}}$  values for these experiments. These results agree qualitatively with those  
582 obtained in previous studies of ozone uptake by metal oxides<sup>25,25,31</sup> and clays<sup>8,29</sup>, and most  
583 likely reflect competition by water vapour for available surface sites for ozone  
584 adsorption<sup>46</sup>.

## 585 **2.2 Ozone uptake by TiO<sub>2</sub> under illuminated conditions**

### 586 **2.2.1 Mechanistic discussion for photoenhanced ozone uptake by TiO<sub>2</sub>**

587 Here, we elaborate on the mechanism proposed by Nicolas et al. for the TiO<sub>2</sub>-catalyzed  
588 uptake of ozone (**R5-R13**)<sup>47</sup>:





598 In this mechanism, absorption of light with energy equal to or greater than the band gap of  
 599  $\text{TiO}_2$  leads to the formation of an electron-hole pair (**R5**) and initiates a chain of reactions  
 600 that reduce ozone and ultimately result in the release of molecular oxygen (**R6–R11**). In  
 601 this proposed mechanism, ozone can be reduced directly by photogenerated electrons (**R6**)  
 602 or indirectly by superoxide radicals produced from the reduction of molecular oxygen by  
 603 photogenerated electrons (**R7–8**). In both cases, the ozonide radical thus produced can  
 604 decompose to yield hydroxyl radicals (**R9–10**), which can in turn lead to further ozone  
 605 decomposition (**R11**). Under our experimental conditions, additional mechanisms for  
 606 hydroxyl radical production include the reaction of photogenerated holes with molecularly  
 607 adsorbed water (**R12**) and/or the hydrated surface of  $\text{TiO}_2$  (**R13**)<sup>48</sup>.

### 608 **2.2.2 Comparison of uptake coefficients to previous literature**

609 Although this work is the first to study the photochemical uptake of ozone by natural Ti-  
 610 containing minerals, several other studies have explored the photochemical uptake of  
 611 ozone by commercial  $\text{TiO}_2$ . Using a coated-wall flow tube, Nicolas et al. obtained a  $\gamma_{\text{BET}}$  value  
 612 of  $(2.6 \pm 0.4) \times 10^{-6}$  for  $\text{TiO}_2/\text{SiO}_2$  mixed films at similar ozone concentrations, RH values,

613 and film masses to those employed in our experiments<sup>47</sup>. This value is much lower than the  
614 value we obtain in this work (*i.e.*  $(1.8 \pm 0.3) \times 10^{-4}$ ). We suggest that this difference may  
615 reflect different approaches used to calculate  $\gamma_{\text{BET}}$ : specifically, because these authors were  
616 interested in the reactivity of the composite film rather than the reactivity of TiO<sub>2</sub> itself,  
617 they reported  $\gamma_{\text{BET}}$  values corrected for the entire surface area of their TiO<sub>2</sub>/SiO<sub>2</sub> mixed  
618 films. By contrast, in this work, because we were interested in obtaining  $\gamma_{\text{BET}}$  for the TiO<sub>2</sub>  
619 fraction, we used the surface area of TiO<sub>2</sub> alone in our calculations (see **Section 1.3.3.1**). In  
620 another study, Chen et al.<sup>31</sup> used an environmental reaction chamber to obtain a  $\gamma_{\text{BET}}$  value  
621 of  $(2.4 \pm 0.1) \times 10^{-7}$ , which is lower than both the values obtained in this study and those  
622 reported by Nicolas et al.<sup>31</sup>. We suggest that this discrepancy likely reflects the use of  
623 significantly higher ozone concentrations in the reaction chamber than in the flow tube  
624 studies, as calculated  $\gamma_{\text{BET}}$  values for surface-mediated reactions display a negative  
625 dependence on gas-phase ozone concentrations<sup>47</sup>. Finally, whereas our commercial TiO<sub>2</sub>  
626 was primarily anatase, the two previous studies used Degussa TiO<sub>2</sub>, which contains both  
627 anatase and rutile. Since studies have shown that the photoreactivity of mixed TiO<sub>2</sub> phases  
628 can differ from that of individual TiO<sub>2</sub> phases<sup>49</sup>, it is also possible that the inherent  
629 photoreactivity of our TiO<sub>2</sub> sample differs from those employed in these previous studies.

### 630 **2.2.3 Influence of relative humidity on ozone uptake**

631 In order to assess the influence of adsorbed water on the photochemical uptake of ozone by  
632 our TiO<sub>2</sub> samples, we performed experiments at several RH values. As illustrated in **Figure**  
633 **S13**, anatase 1 exhibits the highest photoreactivity at RH 10%. These results agree with  
634 those of Nicolas et al., who observed the highest photoreactivity values for TiO<sub>2</sub> with ozone

635 at intermediate RH values<sup>47</sup>. They attributed lower photoreactivities under wet conditions  
636 to competitive adsorption by water vapour, and lower photoreactivities under dry  
637 conditions to reductions in the quantity of adsorbed water available to generate radical  
638 species ((**R12**) in the mechanism described in **Section 2.2.1**) at the illuminated TiO<sub>2</sub>  
639 surface<sup>47,51</sup>. The fact that photochemistry still occurs under dry conditions in both of our  
640 studies implies that the hydrated TiO<sub>2</sub> surface itself may also participate in photochemistry  
641 (**R13**); additional evidence for photochemistry in the absence of water vapour is also  
642 provided by observations of TiO<sub>2</sub>-mediated photochemistry under vacuum conditions (*i.e.*  
643 in a photochemical Knudsen cell)<sup>52</sup>.

644 **Figure S13** also shows that the trends in  $\gamma_{\text{BET}}$  as a function of RH are mineral-dependent  
645 for our sample set: specifically, commercial anatase and anatase 2 display the highest  
646 photoreactivity at RH <0.1%, whereas the photoreactivity of rutile is relatively invariant  
647 with RH. If the two competing mechanisms described by Nicolas et al.<sup>47</sup> are also in  
648 operation for our samples, these discrepancies imply that our samples may differ in their  
649 degree of inherent hydration and/or their ability to retain residual water during our  
650 (relatively mild) sample drying process<sup>53</sup>. These results again highlight the complexities  
651 associated with photochemistry at the surface of natural samples, and that observations  
652 obtained for model systems may not be directly applicable to all environmental substrates.

653 **3** **References**

- 654 (1) Parrish, D. D.; Murphy, P. C.; Albritton, D. L.; Fehsenfeld, F. C. The Measurement of the  
655 Photodissociation Rate of NO<sub>2</sub> in the Atmosphere. *Atmos. Environ.* **1983**, *17* (7), 1365–1379.  
656 [https://doi.org/10.1016/0004-6981\(83\)90411-0](https://doi.org/10.1016/0004-6981(83)90411-0).
- 657 (2) Zafonte, L.; Rieger, P. L.; Holmes, J. R. Nitrogen Dioxide Photolysis in the Los Angeles  
658 Atmosphere. *Environ. Sci. Technol.* **1977**, *11* (5), 483–487.
- 659 (3) Nicolas, M.; Ndour, M.; Ka, O.; D'Anna, B.; George, C. Photochemistry of Atmospheric Dust: Ozone  
660 Decomposition on Illuminated Titanium Dioxide. *Environ. Sci. Technol.* **2009**, *43* (19), 7437–  
661 7442. <https://doi.org/10.1021/es901569d>.
- 662 (4) Pöschl, U.; Rudich, Y.; Ammann, M. Kinetic Model Framework for Aerosol and Cloud Surface  
663 Chemistry and Gas-Particle Interactions–Part 1: General Equations, Parameters, and  
664 Terminology. *Atmos. Chem. Phys.* **2007**, *7* (23), 5989–6023.
- 665 (5) Michel, A. E.; Usher, C. R.; Grassian, V. H. Reactive Uptake of Ozone on Mineral Oxides and  
666 Mineral Dusts. *Atmos. Environ.* **2003**, *37* (23), 3201–3211. [https://doi.org/10.1016/S1352-  
667 2310\(03\)00319-4](https://doi.org/10.1016/S1352-2310(03)00319-4).
- 668 (6) Adams, J. W.; Rodriguez, D.; Cox, R. A. The Uptake of SO<sub>2</sub> on Saharan Dust: A Flow Tube Study.  
669 *Atmos. Chem. Phys.* **2005**, *5* (10), 2679–2689.
- 670 (7) El Zein, A.; Bedjanian, Y. Interaction of NO<sub>2</sub> with TiO<sub>2</sub> Surface under UV Irradiation:  
671 Measurements of the Uptake Coefficient. *Atmos. Chem. Phys.* **2012**, *12* (2), 1013–1020.  
672 <https://doi.org/10.5194/acp-12-1013-2012>.
- 673 (8) Lasne, J.; Romanias, M. N.; Thevenet, F. Ozone Uptake by Clay Dusts under Environmental  
674 Conditions. *ACS Earth Space Chem.* **2018**, *2* (9), 904–914.  
675 <https://doi.org/10.1021/acsearthspacechem.8b00057>.
- 676 (9) Wang, X.; Romanias, M. N.; Thévenet, F.; Rousseau, A. Geocatalytic Uptake of Ozone onto Natural  
677 Mineral Dust. *Catalysts* **2018**, *8* (7), 263. <https://doi.org/10.3390/catal8070263>.
- 678 (10) Knopf, D. A.; Pöschl, U.; Shiraiwa, M. Radial Diffusion and Penetration of Gas Molecules and  
679 Aerosol Particles through Laminar Flow Reactors, Denuders, and Sampling Tubes. *Anal.*  
680 *Chem.* **2015**, *87* (7), 3746–3754. <https://doi.org/10.1021/ac5042395>.
- 681 (11) Sing, K. S. W. Adsorption by Powders and Porous Solids. In *Adsorption by Powders and Porous*  
682 *Solids (Second Edition)*; Rouquerol, F., Rouquerol, J., Sing, K. S. W., Llewellyn, P., Maurin, G.,  
683 Eds.; Academic Press: Oxford, 2014; pp 237–268. [https://doi.org/10.1016/B978-0-08-  
684 097035-6.00007-3](https://doi.org/10.1016/B978-0-08-097035-6.00007-3).
- 685 (12) Donovan, J. J.; Kremser, D.; Fournelle, J. H.; Goemann, K. *Probe for EPMA: Acquisition,*  
686 *Automation and Analysis, Version 11*; Eugene, Oregon, Probe Software, Inc., 2015.
- 687 (13) Donovan, J. J.; Snyder, D. A.; Rivers, M. L. *An Improved Interference Correction for Trace*  
688 *Element Analysis*, 2nd ed.; Microbeam Analysis, 1993.
- 689 (14) Armstrong, J. T. *CITZAF: A Package of Correction Programs for the Quantitative Electron*  
690 *Microbeam X-Ray-Analysis of Thick Polished Materials, Thin-Films, and Particles*; Microbeam  
691 Analysis 4, 1995.
- 692 (15) Jarosewich, E.; Nelen, J. A.; Norberg, J. A. *Reference Samples for Electron Microprobe Analysis*,  
693 4th ed.; Geostand. Geoanal. Res, 1980.
- 694 (16) Gates-Rector, S.; Blanton, T. The Powder Diffraction File: A Quality Materials Characterization  
695 Database. *Powder Diffraction* **2019**, *34* (4), 352–360.  
696 <https://doi.org/10.1017/S0885715619000812>.
- 697 (17) Jade 9.5; Materials Data: Livermore, CA. 2019.
- 698 (18) TOPAS V4: General Profile and Structure Analysis Software for Powder Diffraction Data. -  
699 User's Manual, Bruker AXS, Karlsruhe, Germany. Bruker AXS (2008).
- 700 (19) Pawley, G. S. Unit-Cell Refinement from Powder Diffraction Scans. *J. Appl. Cryst.* **1981**, *14* (6),  
701 357–361. <https://doi.org/10.1107/S0021889881009618>.

- 702 (20) Bruker AXS Inc. *Version, S. A. D. A. B. S.*; Madison WI, USA, 2001.
- 703 (21) Sheldrick, G. M. SHELXTL, Version 6.12. Bruker AXS Inc.: Madison, WI 2001.
- 704 (22) Kortüm, G. *Reflectance Spectroscopy: Principles, Methods, Applications*; Springer Science &  
705 Business Media, 2012.
- 706 (23) Murphy, A. B. Band-Gap Determination from Diffuse Reflectance Measurements of  
707 Semiconductor Films, and Application to Photoelectrochemical Water-Splitting. *Sol. Energy*  
708 *Mater. Sol. Cell* **2007**, *91* (14), 1326–1337. <https://doi.org/10.1016/j.solmat.2007.05.005>.
- 709 (24) Jenkins, R.; Snyder, R. *Introduction to X-Ray Powder Diffraction*; Wiley: New York, 1996.
- 710 (25) Sullivan, R. C.; Thornberry, T.; Abbatt, J. P. D. Ozone Decomposition Kinetics on Alumina:  
711 Effects of Ozone Partial Pressure, Relative Humidity and Repeated Oxidation Cycles. *Atmos.*  
712 *Chem. Phys.* **2004**, *4* (5), 1301–1310. <https://doi.org/10.5194/acp-4-1301-2004>.
- 713 (26) Li, W.; Gibbs, G. V.; Oyama, S. T. Mechanism of Ozone Decomposition on a Manganese Oxide  
714 Catalyst. 1. In Situ Raman Spectroscopy and Ab Initio Molecular Orbital Calculations. *J. Am.*  
715 *Chem. Soc.* **1998**, *120* (35), 9041–9046. <https://doi.org/10.1021/ja981441+>.
- 716 (27) Hanisch, F.; Crowley, J. N. Ozone Decomposition on Saharan Dust: An Experimental  
717 Investigation. *Atmos. Chem. Phys.* **2003**, *3* (1), 119–130. [https://doi.org/10.5194/acp-3-119-](https://doi.org/10.5194/acp-3-119-2003)  
718 2003.
- 719 (28) Michel A. E.; Usher C. R.; Grassian V. H. Heterogeneous and Catalytic Uptake of Ozone on  
720 Mineral Oxides and Dusts: A Knudsen Cell Investigation. *Geophys. Res. Lett.* **2002**, *29* (14),  
721 10–11. <https://doi.org/10.1029/2002GL014896>.
- 722 (29) Coates Fuentes, Z. L.; Kucinski, T. M.; Hinrichs, R. Z. Ozone Decomposition on Kaolinite as a  
723 Function of Monoterpene Exposure and Relative Humidity. *ACS Earth Space Chem.* **2017**.  
724 <https://doi.org/10.1021/acsearthspacechem.7b00107>.
- 725 (30) Mogili, P. K.; Kleiber, P. D.; Young, M. A.; Grassian, V. H. Heterogeneous Uptake of Ozone on  
726 Reactive Components of Mineral Dust Aerosol: An Environmental Aerosol Reaction Chamber  
727 Study. *J. Phys. Chem. A* **2006**, *110* (51), 13799–13807. <https://doi.org/10.1021/jp063620g>.
- 728 (31) Chen, H.; Stanier, C. O.; Young, M. A.; Grassian, V. H. A Kinetic Study of Ozone Decomposition  
729 on Illuminated Oxide Surfaces. *J. Phys. Chem. A* **2011**, *115* (43), 11979–11987.  
730 <https://doi.org/10.1021/jp208164v>.
- 731 (32) Maters, E. C.; Delmelle, P.; Rossi, M. J.; Ayris, P. M.; Bernard, A. Controls on the Surface  
732 Chemical Reactivity of Volcanic Ash Investigated with Probe Gases. *Earth Planet. Sci. Lett.*  
733 **2016**, *450*, 254–262. <https://doi.org/10.1016/j.epsl.2016.06.044>.
- 734 (33) Bulanin, K. M.; Lavalley, J. C.; Tsyganenko, A. A. IR Spectra of Adsorbed Ozone. *Colloids Surf. A*  
735 *Physicochem. Eng. Asp.* **1995**, *101* (2), 153–158. [https://doi.org/10.1016/0927-](https://doi.org/10.1016/0927-7757(95)03130-6)  
736 7757(95)03130-6.
- 737 (34) Lampimäki, M.; Zelenay, V.; Křepelová, A.; Liu, Z.; Chang, R.; Bluhm, H.; Ammann, M. Ozone-  
738 Induced Band Bending on Metal-Oxide Surfaces Studied under Environmental Conditions.  
739 *ChemPhysChem* **2013**, *14* (11), 2419–2425. <https://doi.org/10.1002/cphc.201300418>.
- 740 (35) Li, W.; Oyama, S. T. Mechanism of Ozone Decomposition on a Manganese Oxide Catalyst. 2.  
741 Steady-State and Transient Kinetic Studies. *J. Am. Chem. Soc.* **1998**, *120* (35), 9047–9052.  
742 <https://doi.org/10.1021/ja9814422>.
- 743 (36) Radhakrishnan, R.; Oyama, S. T. Ozone Decomposition over Manganese Oxide Supported on  
744 ZrO<sub>2</sub> and TiO<sub>2</sub>: A Kinetic Study Using in Situ Laser Raman Spectroscopy. *J. Catal.* **2001**, *199*  
745 (2), 282–290. <https://doi.org/10.1006/jcat.2001.3167>.
- 746 (37) Roscoe, J. M.; Abbatt, J. P. D. Diffuse Reflectance FTIR Study of the Interaction of Alumina  
747 Surfaces with Ozone and Water Vapor. *J. Phys. Chem. A* **2005**, *109* (40), 9028–9034.  
748 <https://doi.org/10.1021/jp050766r>.
- 749 (38) Fitzpatrick, R. W.; Chittleborough, D. J. Titanium and Zirconium Minerals. In *Soil Mineralogy*  
750 *with Environmental Applications*; John Wiley & Sons, Ltd, 2002; pp 667–690.  
751 <https://doi.org/10.2136/sssabookser7.c22>.

- 752 (39) Chang, R. Y.-W.; Sullivan, R. C.; Abbatt, J. P. D. Initial Uptake of Ozone on Saharan Dust at  
753 Atmospheric Relative Humidities. *Geophys. Res. Lett.* **2005**, *32* (14), L14815.  
754 <https://doi.org/10.1029/2005GL023317>.
- 755 (40) Kolb, C. E.; Cox, R. A.; Abbatt, J. P. D.; Ammann, M.; Davis, E. J.; Donaldson, D. J.; Garrett, B. C.;  
756 George, C.; Griffiths, P. T.; Hanson, D. R.; Kulmala, M.; McFiggans, G.; Pöschl, U.; Riipinen, I.;  
757 Rossi, M. J.; Rudich, Y.; Wagner, P. E.; Winkler, P. M.; Worsnop, D. R.; O' Dowd, C. D. An  
758 Overview of Current Issues in the Uptake of Atmospheric Trace Gases by Aerosols and  
759 Clouds. *Atmospheric Chemistry and Physics* **2010**, *10* (21), 10561–10605.  
760 <https://doi.org/10.5194/acp-10-10561-2010>.
- 761 (41) Hanisch, F.; Crowley, J. N. Ozone Decomposition on Saharan Dust: An Experimental  
762 Investigation. *Atmos. Chem. Phys.* **2003**, *3* (1), 119–130. [https://doi.org/10.5194/acp-3-119-](https://doi.org/10.5194/acp-3-119-2003)  
763 [2003](https://doi.org/10.5194/acp-3-119-2003).
- 764 (42) Aubin, D. G.; Abbatt, J. P. D. Interaction of NO<sub>2</sub> with Hydrocarbon Soot: Focus on HONO Yield,  
765 Surface Modification, and Mechanism. *J Phys Chem A* **2007**, *111* (28), 6263–6273.  
766 <https://doi.org/10.1021/jp068884h>.
- 767 (43) Wu, L. Impact of Saharan Air Layer on Hurricane Peak Intensity. *Geophys. Res. Lett.* **2007**, *34*  
768 (9). <https://doi.org/10.1029/2007GL029564>.
- 769 (44) Perry, K. D.; Cahill, T. A.; Eldred, R. A.; Dutcher, D. D.; Gill, T. E. Long-Range Transport of North  
770 African Dust to the Eastern United States. *J. Geophys. Res.* **1997**, *102* (D10), 11225–11238.  
771 <https://doi.org/10.1029/97JD00260>.
- 772 (45) Xie, C.; Nishizawa, T.; Sugimoto, N.; Matsui, I.; Wang, Z. Characteristics of Aerosol Optical  
773 Properties in Pollution and Asian Dust Episodes over Beijing, China. *Appl. Opt.* **2008**, *47* (27),  
774 4945–4951. <https://doi.org/10.1364/AO.47.004945>.
- 775 (46) Rubasinghege, G.; H. Grassian, V. Role(s) of Adsorbed Water in the Surface Chemistry of  
776 Environmental Interfaces. *Chem. Commun.* **2013**, *49* (30), 3071–3094.  
777 <https://doi.org/10.1039/C3CC38872G>.
- 778 (47) Nicolas, M.; Ndour, M.; Ka, O.; D'Anna, B.; George, C. Photochemistry of Atmospheric Dust:  
779 Ozone Decomposition on Illuminated Titanium Dioxide. *Environ. Sci. Technol.* **2009**, *43* (19),  
780 7437–7442. <https://doi.org/10.1021/es901569d>.
- 781 (48) Hoffmann, M. R.; Martin, S. T.; Choi, W.; Bahnemann, D. W. Environmental Applications of  
782 Semiconductor Photocatalysis. *Chem. Rev.* **1995**, *95* (1), 69–96.  
783 <https://doi.org/10.1021/cr00033a004>.
- 784 (49) Hurum, D. C.; Agrios, A. G.; Gray, K. A.; Rajh, T.; Thurnauer, M. C. Explaining the Enhanced  
785 Photocatalytic Activity of Degussa P25 Mixed-Phase TiO<sub>2</sub> Using EPR. *J. Phys. Chem. B* **2003**,  
786 *107* (19), 4545–4549. <https://doi.org/10.1021/jp0273934>.
- 787 (50) Bickley, R. I.; Gonzalez-Carreno, T.; Lees, J. S.; Palmisano, L.; Tilley, R. J. D. A Structural  
788 Investigation of Titanium Dioxide Photocatalysts. *J. Solid State Chem.* **1991**, *92* (1), 178–190.  
789 [https://doi.org/10.1016/0022-4596\(91\)90255-G](https://doi.org/10.1016/0022-4596(91)90255-G).
- 790 (51) Folli, A.; Campbell, S. B.; Anderson, J. A.; Macphee, D. E. Role of TiO<sub>2</sub> Surface Hydration on NO  
791 Oxidation Photo-Activity. *J. Photoch. Photobio. A* **2011**, *220* (2), 85–93.  
792 <https://doi.org/10.1016/j.jphotochem.2011.03.017>.
- 793 (52) Styler, S. A.; Donaldson, D. J. Photooxidation of Atmospheric Alcohols on Laboratory Proxies  
794 for Mineral Dust. *Environ. Sci. Technol.* **2011**, *45* (23), 10004–10012.  
795 <https://doi.org/10.1021/es202263q>.
- 796 (53) Brown, G. E.; Henrich, V. E.; Casey, W. H.; Clark, D. L.; Eggleston, C.; Felmy, A.; Goodman, D. W.;  
797 Grätzel, M.; Maciel, G.; McCarthy, M. I.; Neelson, K. H.; Sverjensky, D. A.; Toney, M. F.; Zachara,  
798 J. M. Metal Oxide Surfaces and Their Interactions with Aqueous Solutions and Microbial  
799 Organisms. *Chem. Rev.* **1999**, *99* (1), 77–174. <https://doi.org/10.1021/cr980011z>.

- 800 (54) Ivanov, A. V.; Trakhtenberg, S.; Bertram, A. K.; Gershenson, Y. M.; Molina, M. J. OH, HO<sub>2</sub>, and  
801 Ozone Gaseous Diffusion Coefficients. *J. Phys. Chem. A* **2007**, *111* (9), 1632–1637.  
802 <https://doi.org/10.1021/jp066558w>.  
803 (55) Jennings, S. G. The Mean Free Path in Air. *J. Aerosol Sci* **1988**, *19* (2), 159–166.  
804 [https://doi.org/10.1016/0021-8502\(88\)90219-4](https://doi.org/10.1016/0021-8502(88)90219-4).  
805

806 **4 Supporting figures and tables**

807

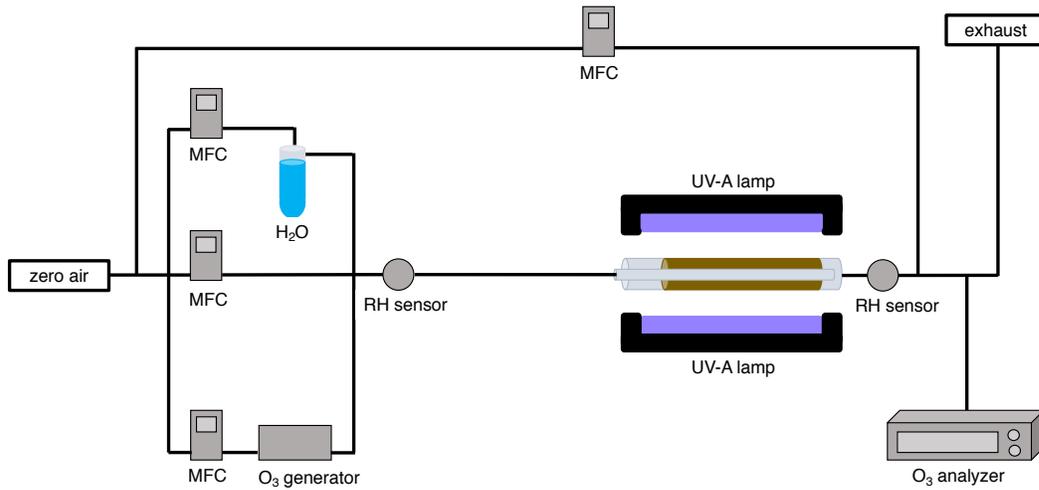
808 **Figure S1**

809 Schematic of the photochemical coated-wall flow tube reactor, with a) the movable injector  
810 pushed in, which prevents the interaction of ozone with mineral samples and b) the  
811 movable injector pulled back, which allows for interaction between ozone and mineral  
812 samples. For clarity, only 2 of the 4 UV-A lamps in the apparatus are shown here.

813

814 a)

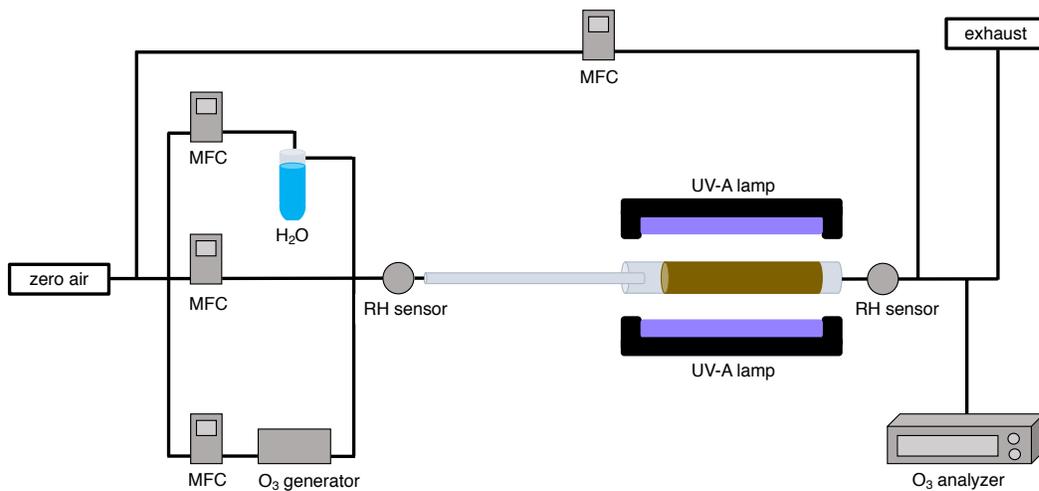
815



816

817

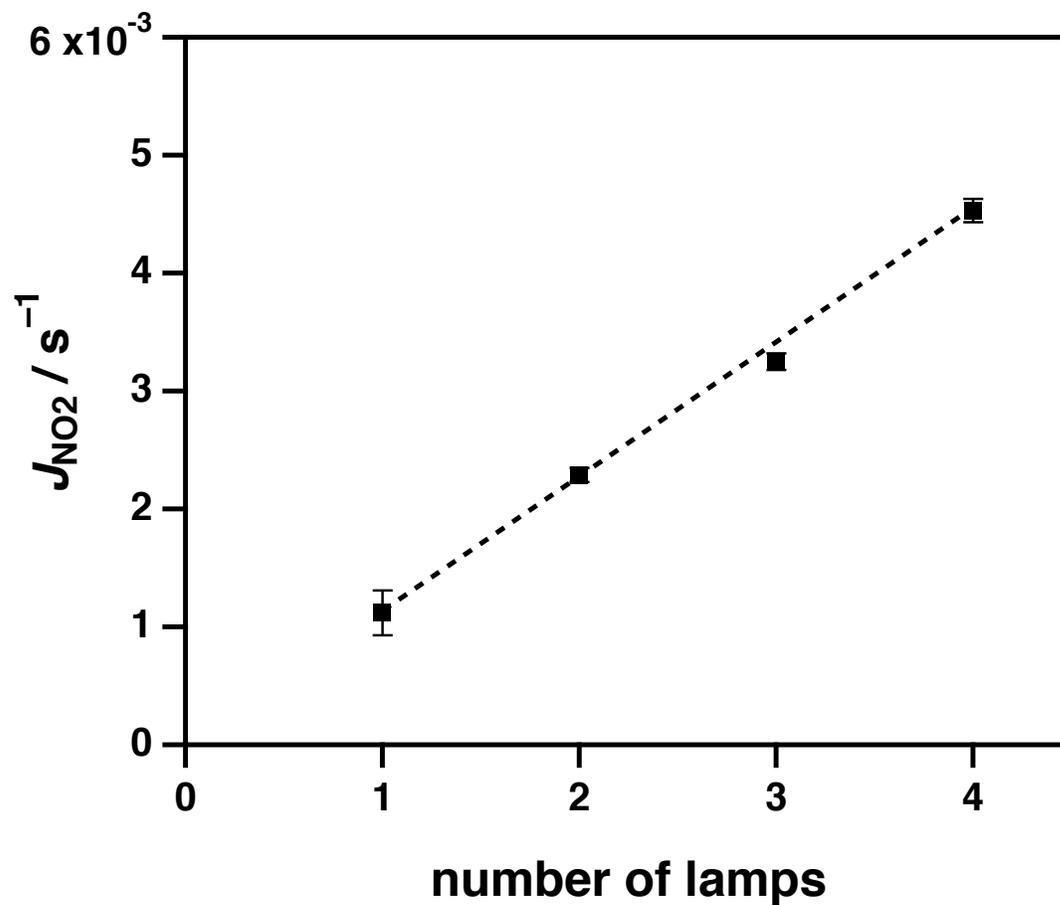
818 b)



819

820

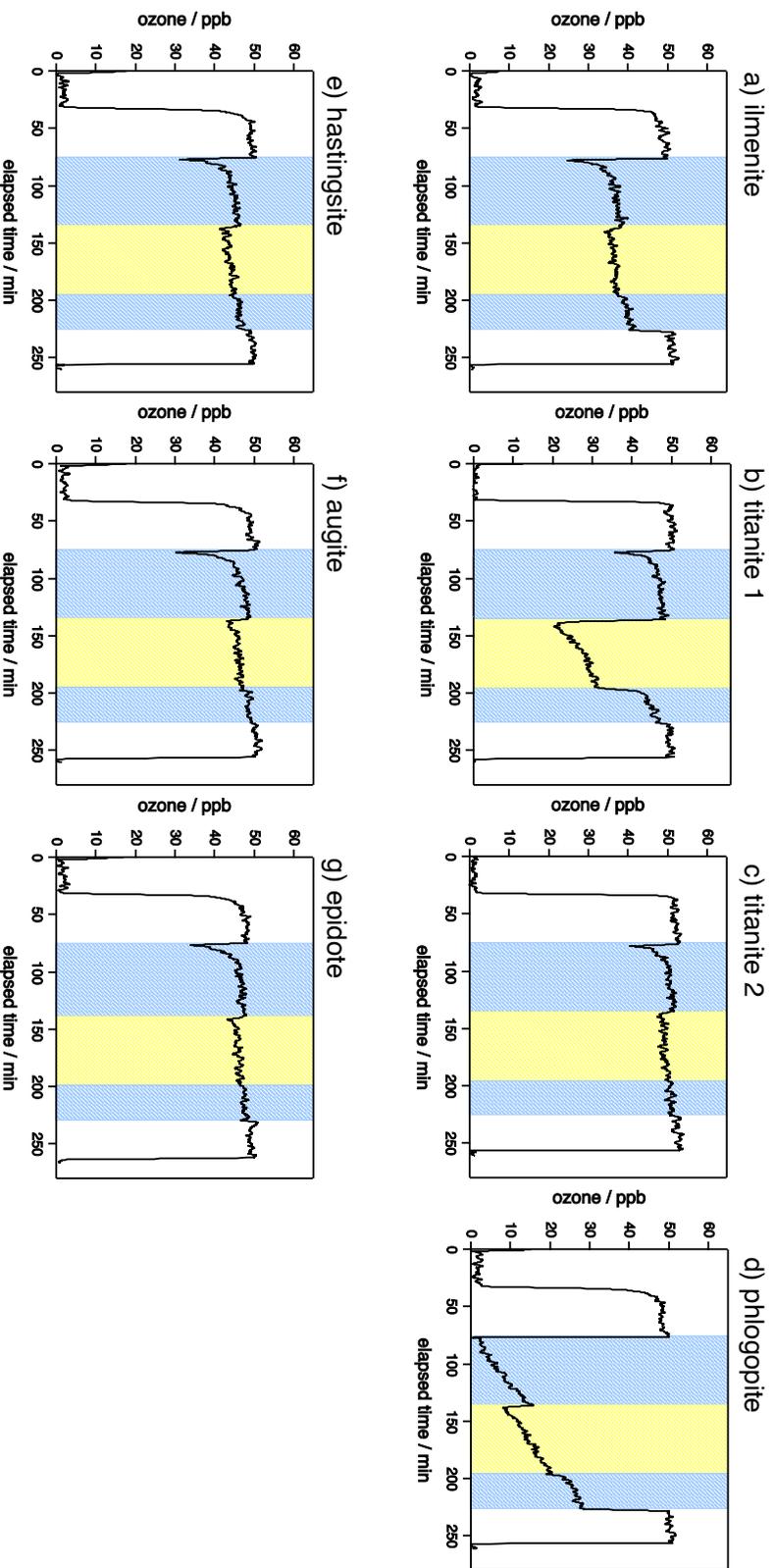
821 **Figure S2**  
822 Photolysis frequencies for  $\text{NO}_2$  ( $J_{\text{NO}_2}$ ) inside the flow tube versus the number of lamps  
823 turned on. Each data point represents the mean of 3 experimental trials, with  $1\sigma$  error  
824 bars; in some cases, error bars are too small to be seen. The dashed line is a linear fit to the  
825 experimental data.  
826



827

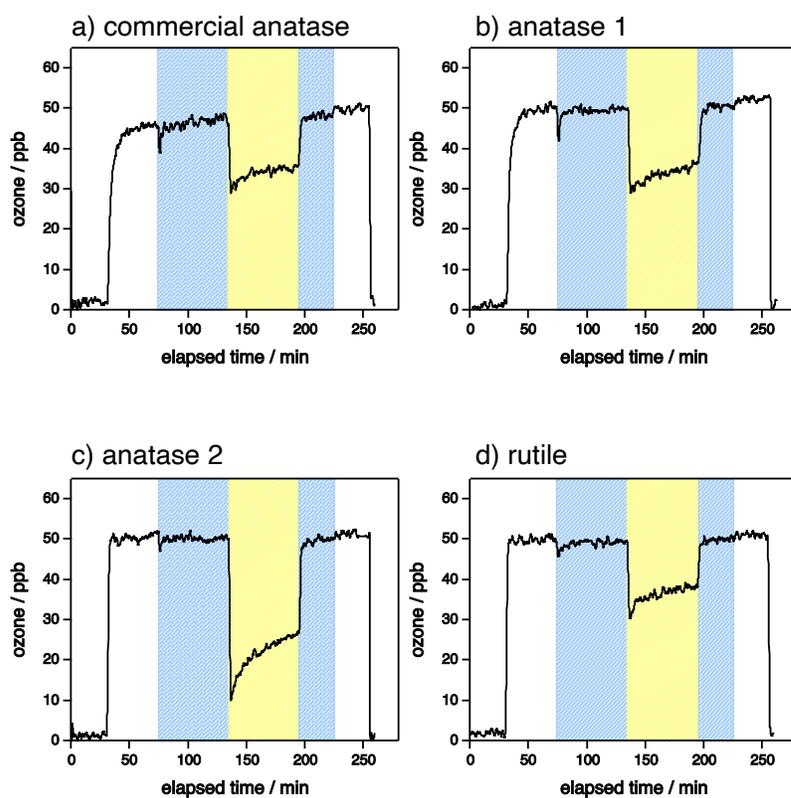
### Figure S3

Representative reaction profiles of Ti-containing minerals with ozone at 25% RH. The unshaded regions of the profile show the time periods in which ozone was not exposed to our Ti-containing mineral. The blue and yellow regions of the profile show the exposure of ozone to our Ti-containing mineral under dark and illuminated conditions, respectively: a) ilmenite, b) titanite 1, c) titanite 2, d) phlogopite, e) hastingsite, f) augite, and g) epidote.



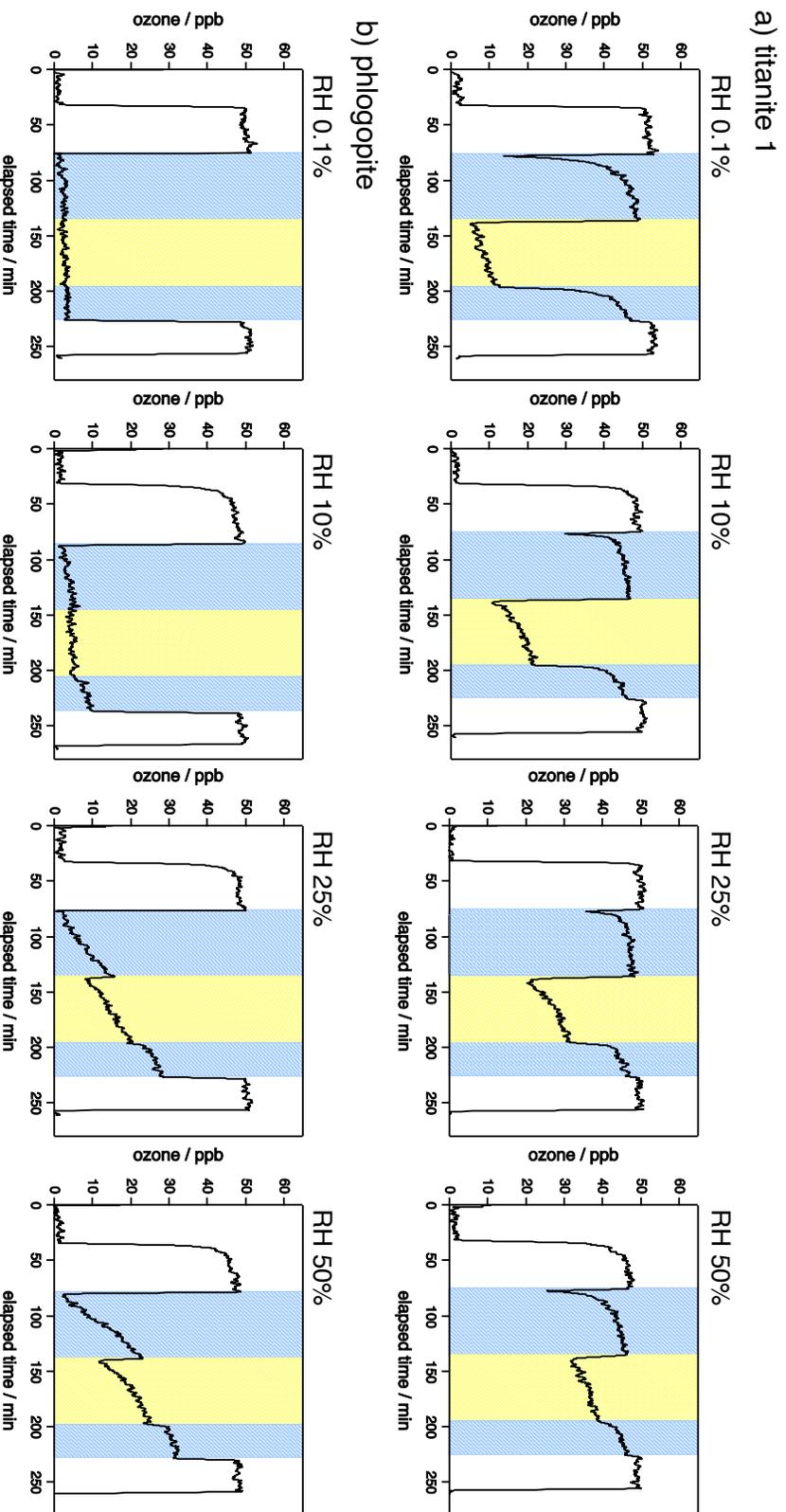
**Figure S4**

Representative reaction profiles of  $\text{TiO}_2/\text{SiO}_2$  mixtures with ozone at 25% RH. The unshaded regions of the profile show time periods in which ozone was not exposed to our Ti-containing mineral. The blue and yellow regions of the profile show the exposure of ozone to our Ti-containing mineral under dark and illuminated conditions, respectively: a) commercial anatase (0.1 wt. %), b) anatase 1 (1 wt. %), c) anatase 2 (20 wt. %), and d) rutile (10 wt. %).



**Figure S5**

Reaction profiles of selected Ti-containing minerals with ozone as a function of RH. The unshaded regions of the profile show time periods in which ozone was not exposed to our Ti-containing mineral. The blue and yellow regions of the profile show the exposure of ozone to our Ti-containing mineral under dark and illuminated conditions, respectively: a) titanite 1 and b) phlogopite.



**Figure S6**

Photographs of natural Ti-containing minerals prior to grinding, with countries of origin listed in parentheses.



**anatase 1**  
(Pakistan)



**anatase 2**  
(Brazil)



**rutile**  
(Brazil)



**ilmenite**  
(United States)



**titanite 1**  
(Canada)



**titanite 2**  
(Brazil)



**phlogopite**  
(Canada)



**hastingsite**  
(United States)



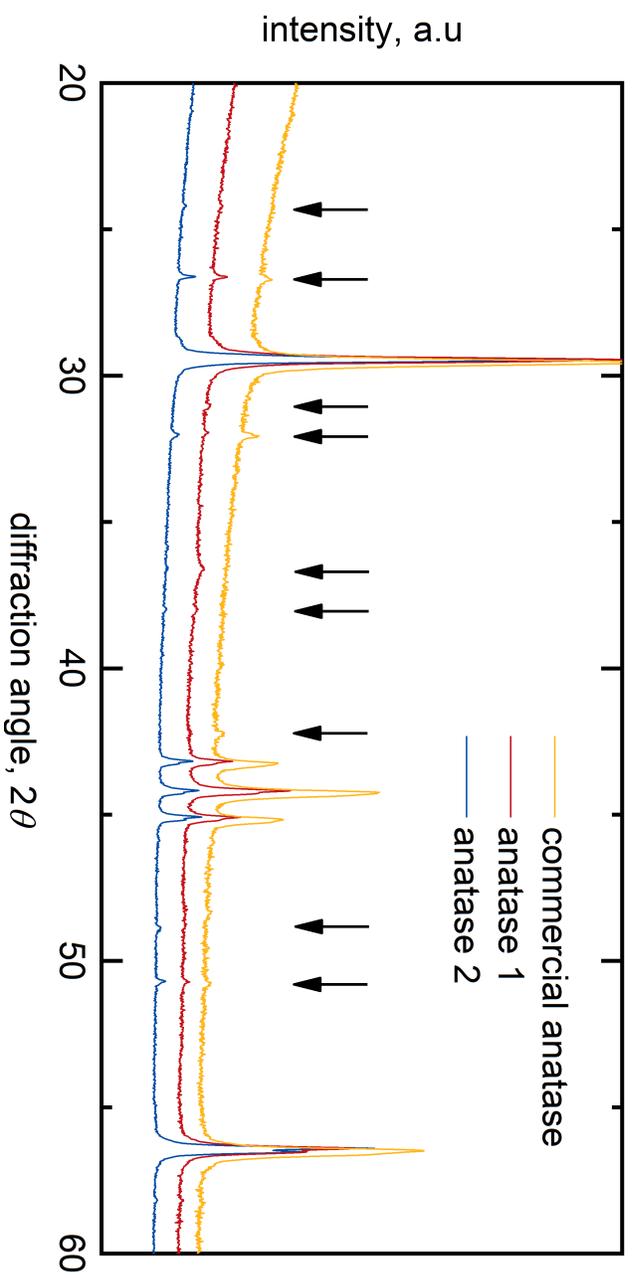
**augite**  
(United States)



**epidote**  
(Norway)

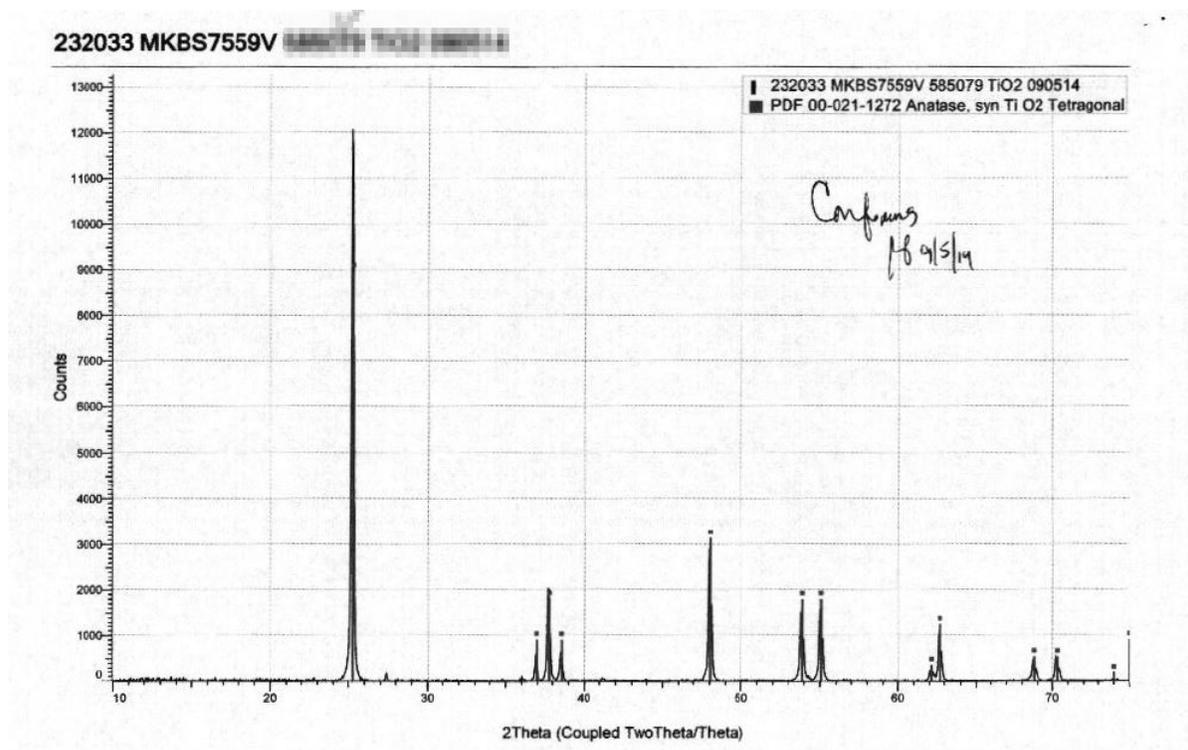
**Figure S7**

Comparison of powder diffraction data of commercial anatase, anatase 1, and anatase 2. Impurity peaks and/or peaks from K $\beta$  radiation are indicated with arrows.



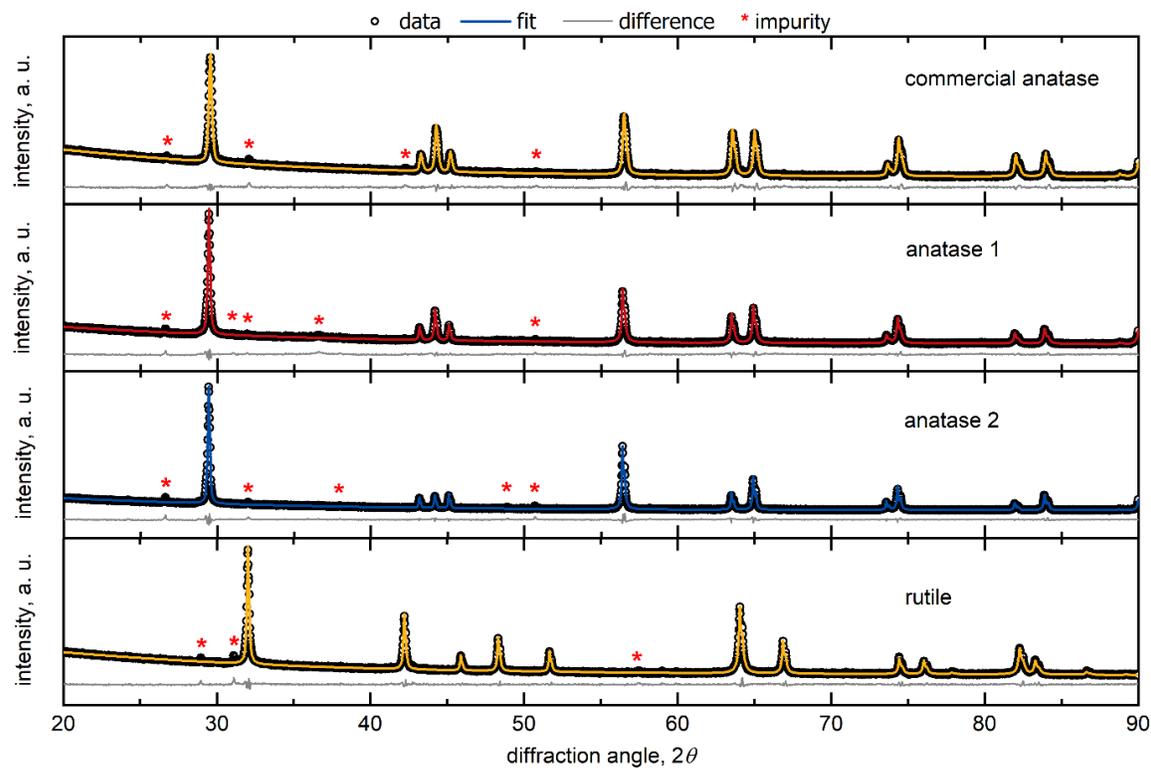
### Figure S8

Anatase purity check from Sigma Aldrich. The pattern is from the same batch as our commercial anatase sample. Cu radiation source was used for this measurement. Minor rutile peak is at  $27.3^\circ$   $2\theta$  angle.



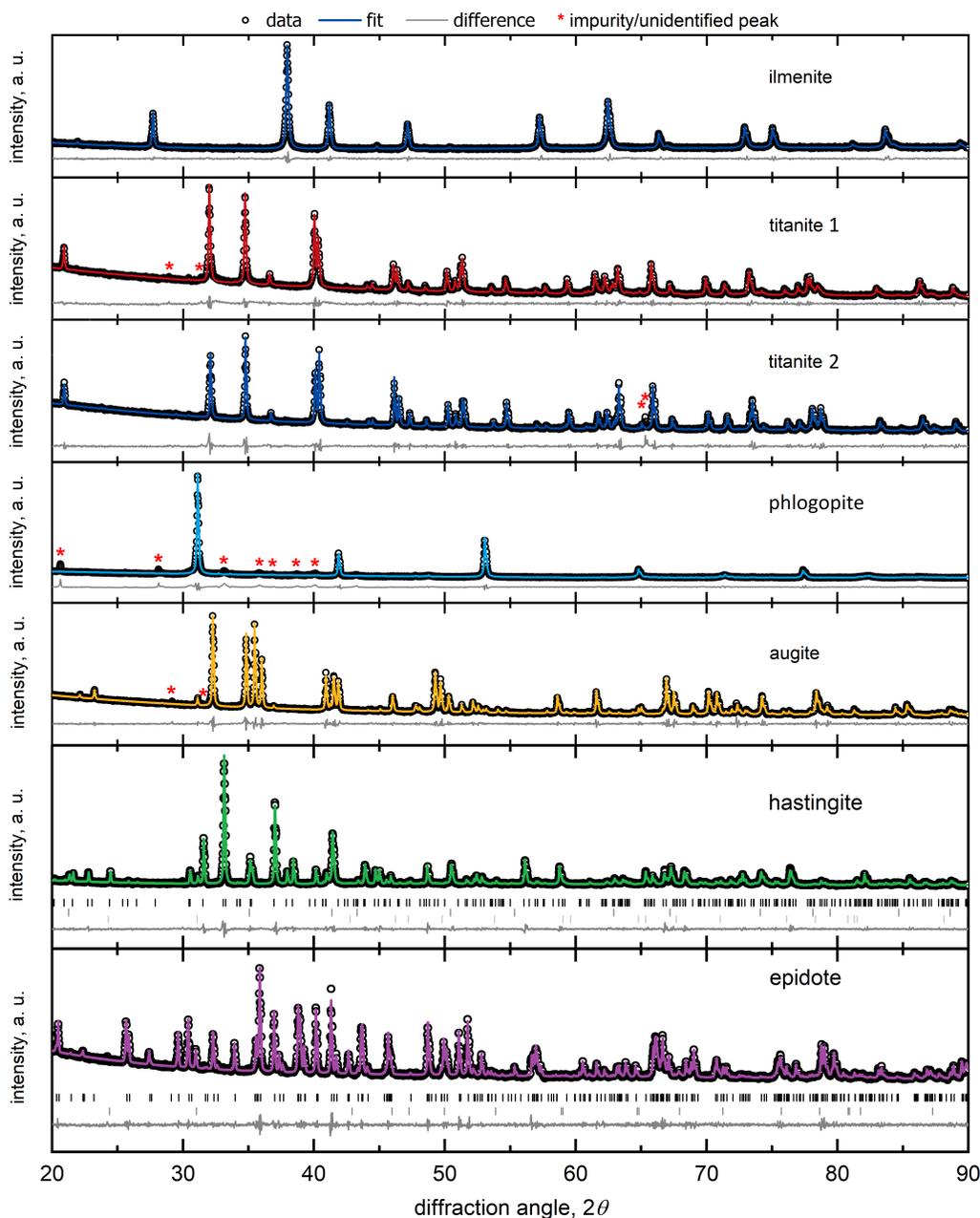
**Figure S9**

Pawley refinement fit for TiO<sub>2</sub> minerals: commercial anatase, anatase 1, anatase 2, and rutile.



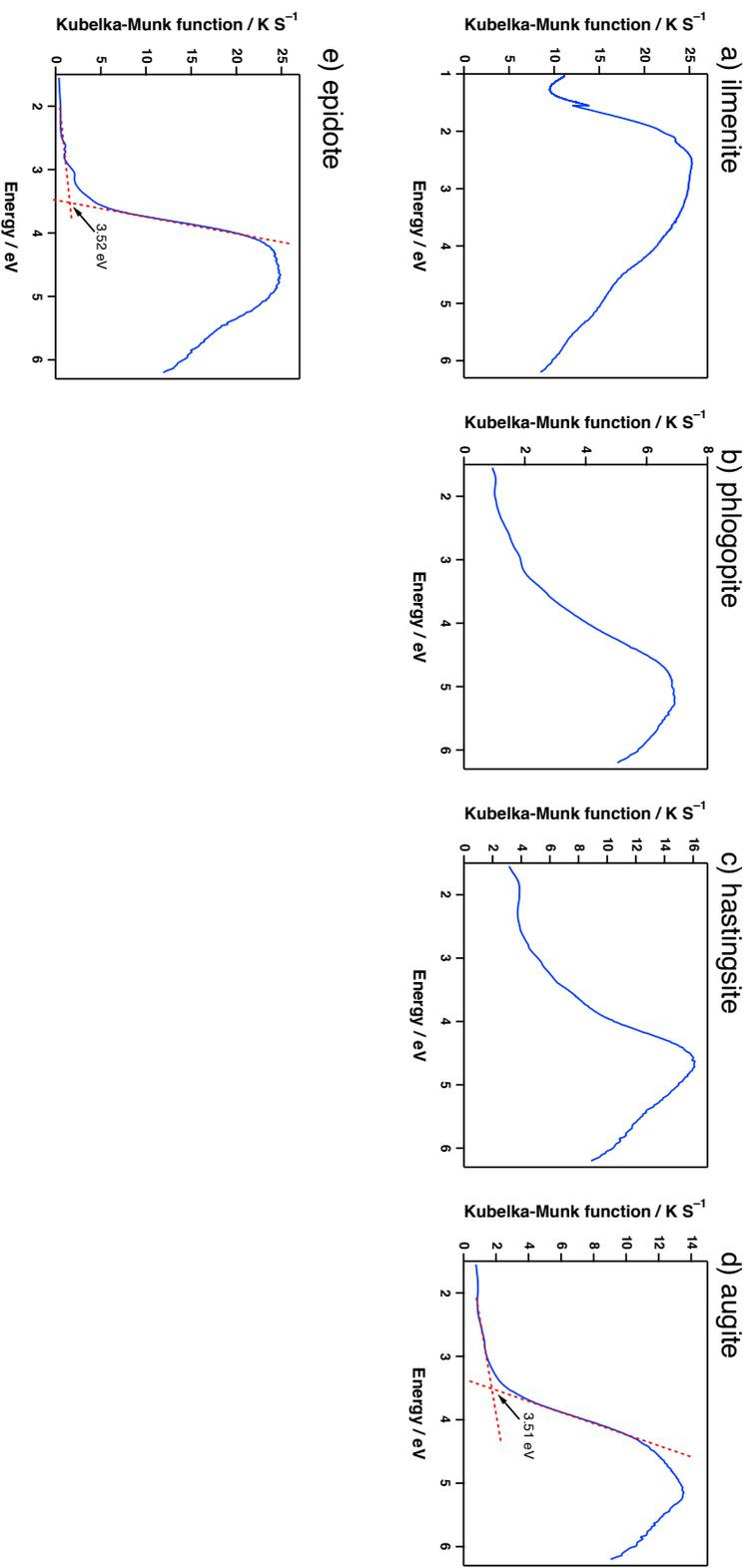
**Figure S10**

Pawley refinement fit for remaining minerals. Ilmenite, titanite 1, titanite 2, phlogopite, and augite are fit with only one phase, whereas the hastingsite sample has diffraction peaks from hastingsite (black ticks), magnetite (dark grey ticks), and quartz (light grey ticks), and the epidote sample has diffraction peaks from epidote (black ticks) and quartz (dark grey ticks).



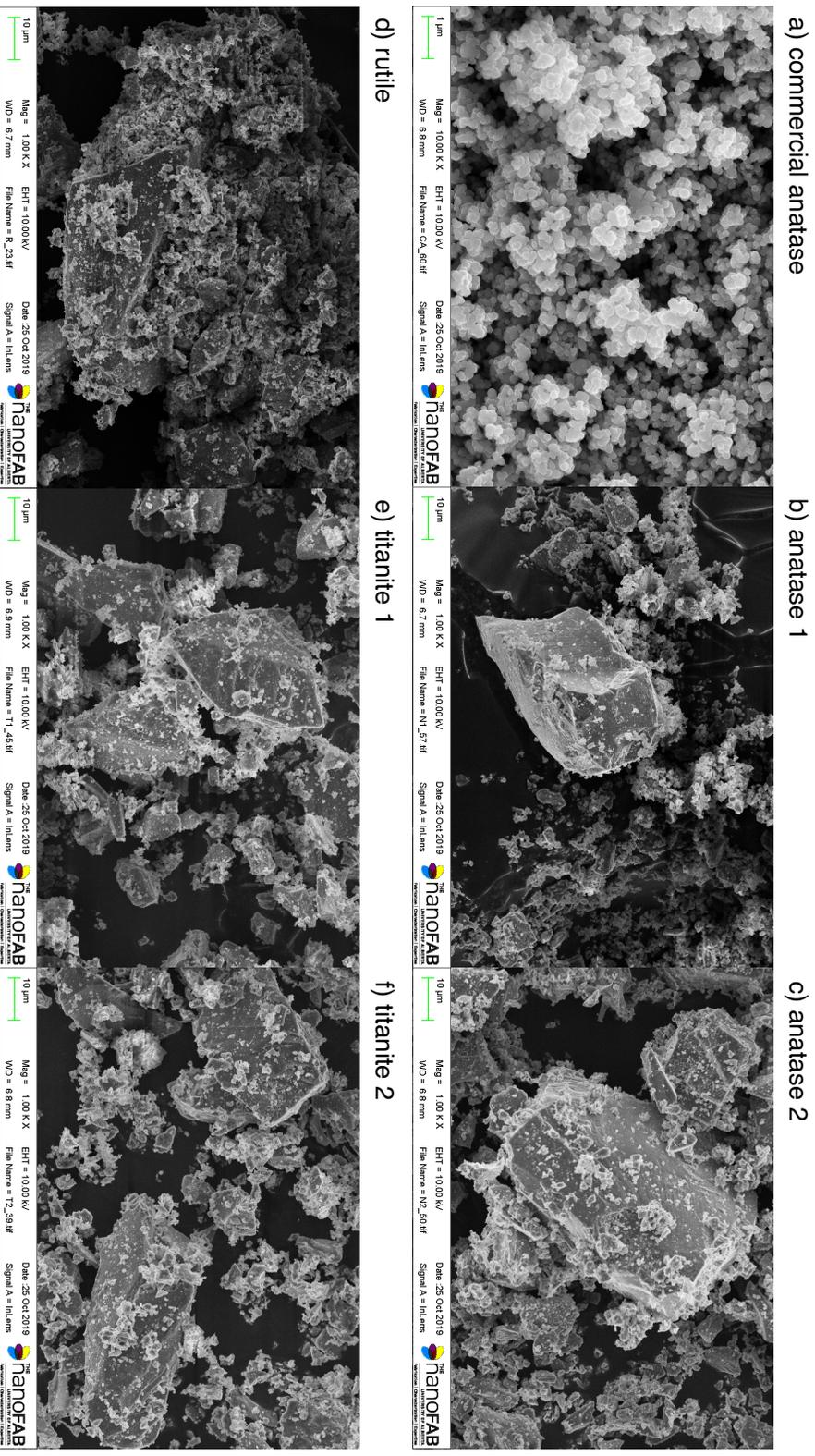
**Figure S11**

Kubelka-Munk plots for Ti-containing minerals, in some cases with associated band gaps: a) ilmenite, b) phlogopite, c) hastingsite, d) augite, and e) epidote.



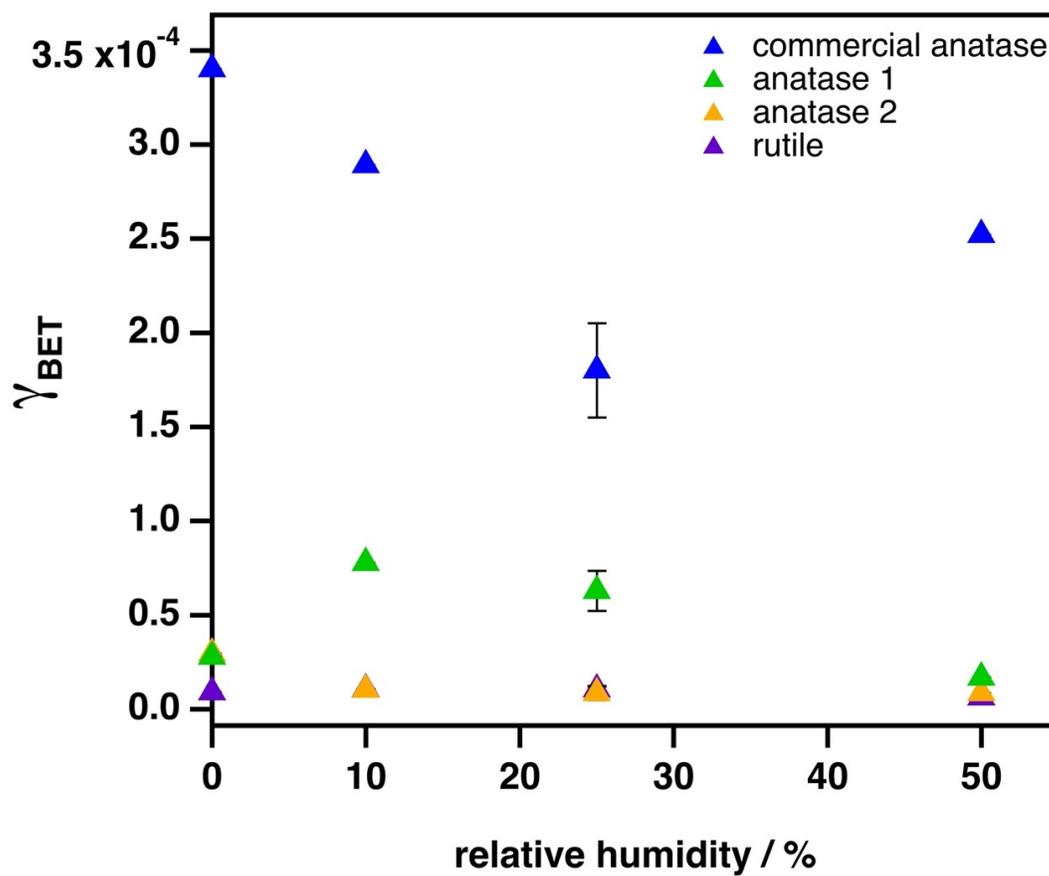
**Figure S12**

Secondary electron images collected by FESEM of: a) commercial anatase, b) anatase 1, c) anatase 2, d) rutile, e) titanite 1, and f) titanite 2.

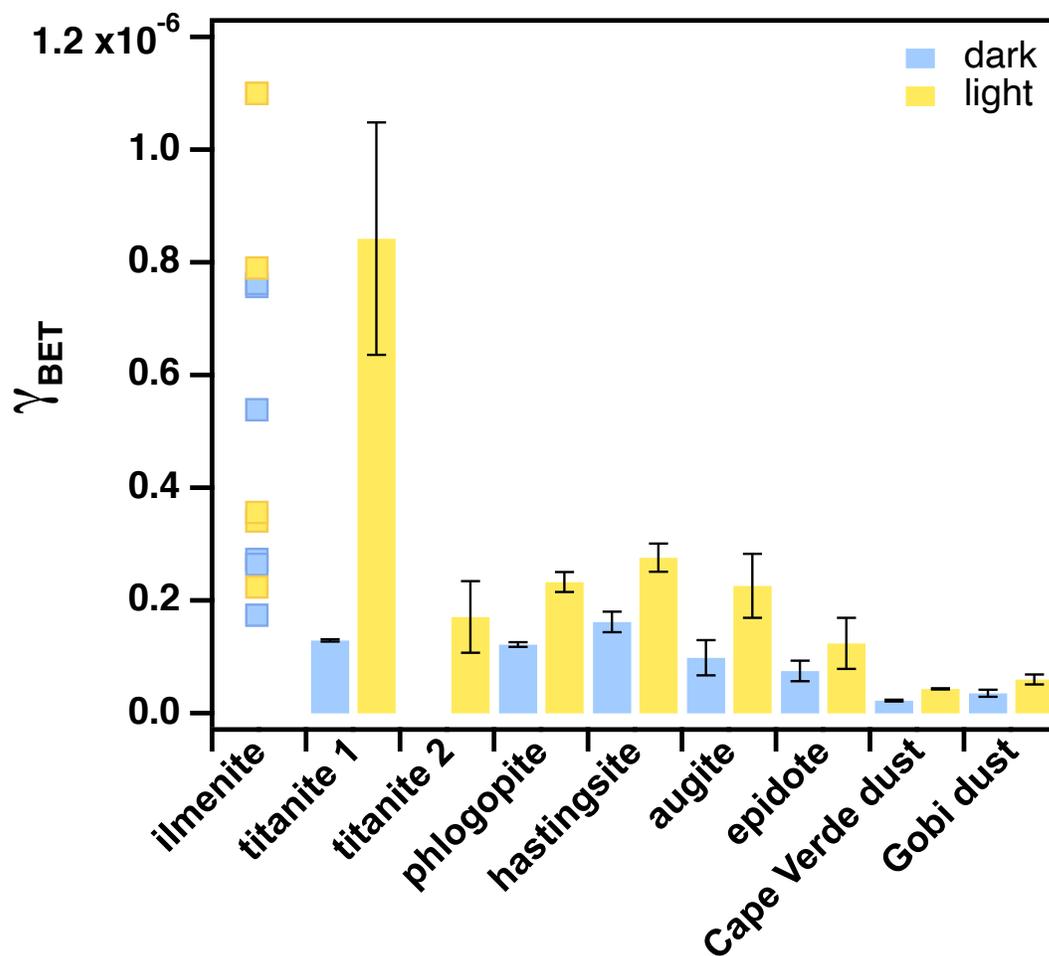


**Figure S13**

Influence of relative humidity on light ozone BET uptake coefficients ( $\gamma_{\text{BET}}$ ) for  $\text{TiO}_2$  minerals. Each data point represents 1 trial, except for RH 25%, in which each data point represents the mean of 3 trials with  $1\sigma$  error bars.



**Figure S14.** BET uptake coefficients ( $\gamma_{\text{BET}}$ ) under dark and illuminated conditions for ozone at RH 25% at the surface of Ti minerals, Ti-bearing minerals, and desert dust samples (Cape Verde dust and Gobi dust). Titanite 2 dark uptake was not significantly larger than the blank experiments (as discussed in main text of the manuscript). Each data point represents the mean of 3 trials with  $1\sigma$  error bars, except for ilmenite, in which each trial ( $n = 6$ ) is displayed individually as this sample appeared to display bimodal reactivity (see Section 2.1).



**Table S1**

Photochemical coated-wall flow tube parameters.

Parameter	Experimental	Equation
Temperature ( $T$ )	296.5 K	
Pressure ( $P$ )	1 atm	
Flow tube reactor diameter ( $D_{reactor}$ )	1.6 cm	
Flow tube reactor length ( $L_{reactor}$ )	34.4 cm	
Pyrex insert tube diameter ( $D_{tube}$ )	1.05 cm	
Pyrex insert tube length ( $L_{tube}$ )	20 cm	
Flow reactor cross-sectional area ( $A$ )	0.865 cm <sup>2</sup>	$A = \pi r^2$
Volumetric flow ( $F$ )	0.440 L min <sup>-1</sup>	
Linear velocity ( $v$ )	8.47 cm s <sup>-1</sup>	$v = \frac{F}{A}$
Length of coated tube ( $L$ )	14.0 cm	
Residence time ( $t$ )	1.6 s	$t = \frac{v}{L}$
Reynolds number ( $Re$ )*	59.3	$Re = \frac{\rho \times D_{tube} \times v}{\eta}$
Length to laminar flow ( $l$ ) <sup>10</sup>	2.18 cm	$l = 0.035 \times Re \times D_{tube}$
Mean molecular velocity ( $\omega$ )**	361.6 m s <sup>-1</sup>	$\omega = \sqrt{\frac{8RT}{\pi M}}$
Ozone diffusion coefficient ( $D$ ) <sup>54</sup>	0.13 cm <sup>2</sup> s <sup>-2</sup>	
Mean free path ( $\lambda$ )	105.7 nm	$\lambda = \frac{D}{\omega}$
Knudsen number ( $Kn$ )	$2.00 \times 10^{-5}$	$Kn = \frac{2 \times \lambda}{D_{tube}}$
Dimensionless axial distance ( $z^*$ ) <sup>10</sup>	0.38	$z^* = z \frac{\pi D}{2F}$

		$z = \text{axial distance of tube (L)}$
Effective Sherwood number ( $N_{Shw}^{eff}$ ) <sup>10</sup>	3.97	$N_{Shw}^{eff} = 3.6568 + \frac{A}{(z^* + B)} 55$ $A = 0.0978 \text{ and } B = 0.0154$

\*Re calculations require density ( $\rho$ ) and viscosity ( $\eta$ ) of air, which are taken to be  $1.2 \text{ kg m}^{-3}$  and  $1.8 \times 10^{-5} \text{ Pa s}$ , respectively at  $296.5 \text{ K}$ <sup>55</sup>

\*\* $\omega$  calculations require the ideal gas constant (R) and the molar mass of ozone (M), which are  $8.3145 \text{ J K}^{-1} \text{ mol}^{-1}$  and  $48 \times 10^{-3} \text{ kg mol}^{-1}$

**Table S2**

TiO<sub>2</sub>/SiO<sub>2</sub> mixtures prepared for each TiO<sub>2</sub> sample.

<b>Sample</b>	<b>Mass of SiO<sub>2</sub> (mg)</b>	<b>Mass of TiO<sub>2</sub> (mg)</b>	<b>TiO<sub>2</sub> wt. %</b>
Commercial anatase	899.2	0.87	0.1
Anatase 1	891.0	9.00	1.0
Anatase 2	720	180.08	20.0
Rutile	809.7	90.37	10.0

**Table S3**

Summary of mineral sample pretreatments.

Sample	Dremel®	SUPER IRON out® (0.04 g mL <sup>-1</sup> )	HCl (1 M)
Anatase 1	✓		
Anatase 2			
Rutile			
Ilmenite			
Titanite 1	✓		✓
Titanite 2			
Phlogopite	✓	✓	✓
Hastingsite		✓	
Augite		✓	
Epidote	✓	✓	

**Table S4**Specific surface areas (BET;  $\text{m}^2 \text{g}^{-1}$ ) of each mineral sample under study.

<b>Sample</b>	<b>BET surface area (<math>\text{m}^2 \text{g}^{-1}</math>)</b>
Commercial anatase	9.6791
Anatase 1	3.7258
Anatase 2	1.5205
Rutile	1.9323
Ilmenite	2.0914
Titanite 1	1.9616
Titanite 2	1.0077
Phlogopite	18.7799
Hastingsite	1.7424
Augite	1.6235
Epidote	1.3855
Silicon dioxide	94.5500

**Table S5**

Elemental composition of anatase 1 as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.09	0.00	99.63	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	0.00	0.00	99.75
	0.07	0.00	100.12	0.00	0.00	0.08	0.00	0.00	0.00	0.00	0.00	0.00	0.00	100.27
	0.08	0.00	99.19	0.00	0.00	0.10	0.00	0.09	0.00	0.00	0.00	0.00	0.00	99.46
	0.07	0.00	99.70	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	99.77
	0.12	0.02	99.51	0.00	0.01	0.04	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.73
	0.00	0.00	99.30	0.00	0.02	0.06	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.41
	0.00	0.00	99.14	0.00	0.00	0.11	0.00	0.00	0.00	0.00	0.01	0.00	0.00	99.26
	0.10	0.00	99.14	0.00	0.00	0.10	0.00	0.05	0.00	0.00	0.00	0.00	0.00	99.39
<b>Average</b>	0.07	0.00	99.47	0.00	0.00	0.07	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.63
<b>Std Dev.</b>	0.04	0.01	0.34	0.00	0.01	0.04	0.00	0.03	0.00	0.00	0.00	0.00	0.00	0.32
<b>Min</b>	0.00	0.00	99.14	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	99.26
<b>Max</b>	0.12	0.02	100.12	0.00	0.02	0.11	0.00	0.09	0.00	0.00	0.01	0.00	0.00	100.27

**Table S6**  
 Elemental composition of anatase 2 as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	0.00	100.08	0.00	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	100.11
	0.00	0.00	100.02	0.00	0.00	0.00	0.10	0.00	0.00	0.00	0.00	0.00	100.12
	0.00	0.00	99.89	0.03	0.00	0.00	0.09	0.00	0.00	0.00	0.00	0.00	100.01
	0.00	0.00	99.80	0.00	0.00	0.00	0.06	0.00	0.00	0.00	0.00	0.00	99.86
	0.00	0.00	99.68	0.03	0.00	0.00	0.11	0.00	0.00	0.00	0.00	0.00	99.82
	0.00	0.00	99.67	0.00	0.03	0.00	0.05	0.00	0.00	0.00	0.00	0.00	99.75
	0.00	0.00	100.20	0.00	0.00	0.00	0.05	0.00	0.00	0.00	0.00	0.00	100.25
	0.00	0.00	99.81	0.00	0.00	0.00	0.08	0.00	0.00	0.00	0.00	0.00	99.89
	0.00	0.00	99.63	0.00	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.66
	0.00	0.00	99.78	0.00	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.81
	0.04	0.00	99.81	0.00	0.00	0.00	0.06	0.00	0.00	0.00	0.00	0.00	99.91
	0.00	0.00	99.89	0.00	0.00	0.00	0.08	0.00	0.00	0.00	0.00	0.00	99.97
	0.00	0.00	99.79	0.00	0.00	0.00	0.08	0.00	0.00	0.00	0.00	0.00	99.87
	0.00	0.00	99.78	0.00	0.03	0.00	0.08	0.00	0.00	0.00	0.00	0.01	99.90
	0.00	0.00	99.72	0.00	0.00	0.00	0.07	0.00	0.00	0.00	0.00	0.00	99.79
	0.00	0.00	99.79	0.00	0.02	0.00	0.06	0.00	0.00	0.00	0.00	0.00	99.87
<b>Average</b>	0.00	0.00	99.83	0.00	0.01	0.00	0.07	0.00	0.00	0.00	0.00	0.00	99.91
<b>Std Dev.</b>	0.01	0.00	0.15	0.01	0.01	0.00	0.02	0.00	0.00	0.00	0.00	0.00	0.15
<b>Min</b>	0.00	0.00	99.63	0.00	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	99.66
<b>Max</b>	0.04	0.00	100.20	0.03	0.03	0.00	0.11	0.00	0.00	0.00	0.00	0.01	100.25

**Table S7**

Elemental composition of rutile as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.49	0.00	98.73	0.00	0.04	0.00	0.00	0.68	0.00	0.00	0.00	0.00	0.00	99.94
	0.50	0.00	98.74	0.00	0.04	0.05	0.00	0.67	0.00	0.00	0.00	0.00	0.00	100.00
	0.54	0.01	98.72	0.00	0.03	0.12	0.00	0.67	0.00	0.00	0.00	0.00	0.00	100.09
	0.48	0.00	98.63	0.00	0.03	0.19	0.00	0.66	0.00	0.00	0.00	0.00	0.00	99.99
	0.60	0.02	98.74	0.00	0.02	0.12	0.00	0.66	0.00	0.00	0.00	0.00	0.00	100.16
	0.52	0.00	98.93	0.00	0.03	0.14	0.00	0.68	0.00	0.00	0.00	0.00	0.00	100.30
	0.49	0.00	98.90	0.00	0.04	0.13	0.00	0.68	0.00	0.00	0.00	0.00	0.00	100.24
	0.51	0.00	98.86	0.00	0.02	0.14	0.00	0.64	0.00	0.00	0.00	0.00	0.00	100.17
<b>Average</b>	0.52	0.00	98.78	0.00	0.03	0.11	0.00	0.67	0.00	0.00	0.00	0.00	0.00	100.11
<b>Std Dev.</b>	0.04	0.01	0.10	0.00	0.01	0.06	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.13
<b>Min</b>	0.48	0.00	98.63	0.00	0.02	0.00	0.00	0.64	0.00	0.00	0.00	0.00	0.00	99.94
<b>Max</b>	0.60	0.02	98.93	0.00	0.04	0.19	0.00	0.68	0.00	0.00	0.00	0.00	0.00	100.30

**Table S8**  
Elemental composition of ilmenite as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero. We note that the impurity in the ilmenite sample was not used in the overall average for the elemental composition.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	0.02	50.01	0.04	0.00	0.00	0.00	46.09	0.66	0.00	0.00	0.00	0.00	96.82
	0.00	0.00	50.73	0.00	0.00	0.00	0.00	46.03	0.71	0.00	0.00	0.00	0.00	97.47
	0.00	0.00	50.94	0.04	0.00	0.00	0.00	45.86	0.69	0.00	0.00	0.00	0.00	97.53
	0.00	0.01	50.10	0.03	0.00	0.00	0.00	46.43	0.69	0.00	0.00	0.00	0.00	97.26
	0.47	0.00	51.44	0.00	0.00	0.00	0.00	44.53	0.71	0.00	0.00	0.00	0.00	97.15
	0.00	0.00	50.49	0.03	0.00	0.00	0.00	45.87	0.71	0.00	0.00	0.00	0.00	97.1
	4.48	0.00	53.50	0.00	0.02	0.06	0.00	38.96	0.60	0.03	0.00	0.00	0.00	97.65
<b>Average</b>	0.71	0.00	51.03	0.02	0.00	0.01	0.00	44.82	0.68	0.00	0.00	0.00	0.00	97.28
<b>Std Dev.</b>	1.67	0.01	1.19	0.02	0.01	0.02	0.00	2.65	0.04	0.01	0.00	0.00	0.00	0.29
<b>Min</b>	0.00	0.00	50.01	0.00	0.00	0.00	0.00	38.96	0.60	0.00	0.00	0.00	0.00	96.82
<b>Max</b>	4.48	0.02	53.50	0.04	0.02	0.06	0.00	46.43	0.71	0.03	0.00	0.00	0.00	97.65
<b>Impurity</b>	0.00	3.47	14.82	0.00	1.53	0.04	0.00	66.95	0.39	0.14	0.00	0.00	0.00	87.42

**Table S9**  
 Elemental composition of titanite 1 as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	1.14	30.35	34.98	0.00	1.23	0.00	0.00	2.33	0.08	0.06	27.22	0.30	0.00	97.69
	1.08	30.50	34.98	0.00	1.20	0.04	0.00	2.29	0.08	0.04	27.34	0.32	0.00	97.87
	1.13	30.12	34.72	0.00	1.23	0.06	0.00	2.30	0.09	0.05	27.22	0.32	0.00	97.24
	1.16	30.23	34.86	0.00	1.19	0.04	0.00	2.31	0.09	0.08	27.08	0.29	0.00	97.33
	1.11	30.04	34.70	0.00	1.17	0.06	0.00	2.29	0.08	0.06	27.14	0.33	0.00	96.98
	0.93	30.23	34.39	0.00	1.27	0.00	0.00	2.38	0.08	0.08	27.24	0.30	0.00	96.90
	0.91	30.10	34.51	0.00	1.33	0.06	0.00	2.42	0.08	0.08	27.12	0.29	0.00	96.90
	0.98	30.32	34.59	0.00	1.29	0.00	0.00	2.40	0.08	0.07	27.26	0.29	0.00	97.28
<b>Average</b>	1.06	30.24	34.72	0.00	1.24	0.03	0.00	2.34	0.08	0.07	27.20	0.31	0.00	97.27
<b>Std Dev.</b>	0.10	0.15	0.22	0.00	0.05	0.03	0.00	0.05	0.00	0.02	0.08	0.02	0.00	0.36
<b>Min</b>	0.91	30.04	34.39	0.00	1.17	0.00	0.00	2.29	0.08	0.04	27.08	0.29	0.00	96.90
<b>Max</b>	1.16	30.50	34.98	0.00	1.33	0.06	0.00	2.42	0.09	0.08	27.34	0.33	0.00	97.87

**Table S10**  
Elemental composition of titanite 2 as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	30.81	38.44	0.00	0.86	0.00	0.57	0.00	0.03	28.25	0.00	0.00	98.96
	0.00	30.78	38.68	0.00	0.82	0.00	0.53	0.00	0.03	28.22	0.00	0.00	99.06
	0.00	30.72	38.16	0.00	0.86	0.00	0.57	0.00	0.04	28.30	0.12	0.00	98.77
	0.00	30.78	38.17	0.00	0.86	0.00	0.57	0.00	0.02	28.30	0.00	0.00	98.70
	0.00	30.74	38.26	0.00	0.90	0.00	0.57	0.00	0.02	28.33	0.00	0.00	98.82
	0.00	30.76	38.33	0.00	0.88	0.00	0.57	0.03	0.03	28.35	0.00	0.00	98.95
	0.00	30.75	38.42	0.00	0.96	0.00	0.62	0.00	0.02	28.31	0.00	0.00	99.08
	0.00	30.75	38.64	0.00	0.83	0.00	0.52	0.00	0.03	28.42	0.00	0.00	99.19
	0.00	30.70	38.76	0.00	0.71	0.00	0.50	0.02	0.04	28.26	0.00	0.00	98.99
	0.00	30.77	38.46	0.00	0.90	0.00	0.57	0.02	0.03	28.29	0.00	0.00	99.04
	0.00	30.77	38.31	0.00	0.88	0.00	0.58	0.02	0.03	28.28	0.03	0.00	98.90
	0.00	30.85	38.44	0.00	0.90	0.00	0.58	0.00	0.02	28.39	0.00	0.00	99.18
	0.00	30.76	38.51	0.00	0.85	0.00	0.61	0.00	0.03	28.27	0.00	0.00	99.03
	0.00	30.78	38.51	0.00	0.87	0.00	0.58	0.00	0.03	28.37	0.00	0.00	99.14
	0.00	30.79	38.64	0.00	0.86	0.00	0.59	0.02	0.02	28.29	0.00	0.00	99.21
	0.00	30.73	38.83	0.00	0.70	0.00	0.49	0.00	0.02	28.38	0.00	0.00	99.15
<b>Average</b>	0.00	30.77	38.47	0.00	0.85	0.00	0.56	0.01	0.03	28.31	0.01	0.00	99.01
<b>Std Dev.</b>	0.00	0.04	0.20	0.00	0.07	0.00	0.04	0.01	0.01	0.06	0.03	0.00	0.15
<b>Min</b>	0.00	30.70	38.16	0.00	0.70	0.00	0.49	0.00	0.02	28.22	0.00	0.00	98.70
<b>Max</b>	0.00	30.85	38.83	0.00	0.96	0.00	0.62	0.03	0.04	28.42	0.12	0.00	99.21

**Table S11**  
 Elemental composition of phlogopite as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	42.82	0.92	0.08	11.08	0.00	0.00	6.66	0.13	22.43	0.00	0.43	10.07	94.62
	0.00	42.80	0.98	0.08	11.02	0.00	0.00	6.60	0.13	22.23	0.00	0.18	10.51	94.53
	0.00	42.85	1.05	0.07	11.20	0.00	0.00	6.80	0.13	22.05	0.00	0.25	10.31	94.71
	0.00	42.53	1.02	0.09	11.19	0.00	0.00	6.79	0.12	22.26	0.00	0.34	10.17	94.51
	0.00	42.83	1.04	0.07	11.10	0.00	0.00	6.68	0.12	22.13	0.00	0.25	10.39	94.61
	0.00	42.68	1.05	0.08	10.95	0.00	0.00	6.79	0.12	22.05	0.00	0.33	10.25	94.30
	0.00	42.23	1.07	0.06	11.17	0.00	0.00	6.88	0.14	22.15	0.00	0.36	10.17	94.23
<b>Average</b>	0.00	42.68	1.02	0.08	11.10	0.00	0.00	6.74	0.13	22.19	0.00	0.31	10.27	94.50
<b>Std Dev.</b>	0.00	0.23	0.05	0.01	0.09	0.00	0.00	0.10	0.01	0.13	0.00	0.08	0.15	0.18
<b>Min</b>	0.00	42.23	0.92	0.06	10.95	0.00	0.00	6.60	0.12	22.05	0.00	0.18	10.07	94.23
<b>Max</b>	0.00	42.85	1.07	0.09	11.20	0.00	0.00	6.88	0.14	22.43	0.00	0.43	10.51	94.71

**Table S12**

Elemental composition of hastingsite as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	40.95	0.80	0.00	10.59	0.00	0.00	21.01	0.21	8.76	11.17	1.68	2.07	97.24
	0.00	40.79	0.77	0.00	10.58	0.00	0.00	20.96	0.23	8.66	11.14	1.63	2.02	96.78
	0.00	40.86	0.76	0.03	10.42	0.00	0.00	20.96	0.20	8.64	11.16	1.63	2.02	96.68
	0.00	40.72	0.78	0.03	10.43	0.00	0.00	20.98	0.20	8.59	11.08	1.63	2.02	96.46
	0.00	40.58	0.76	0.00	10.53	0.00	0.00	21.02	0.21	8.73	11.14	1.67	2.02	96.66
	0.00	40.63	0.84	0.03	10.46	0.00	0.00	20.91	0.20	8.62	11.12	1.64	2.07	96.52
	0.00	40.75	0.76	0.00	10.37	0.00	0.00	21.00	0.19	8.66	11.11	1.69	2.03	96.56
	0.00	40.79	0.81	0.00	10.44	0.00	0.00	21.00	0.22	8.68	11.11	1.64	2.02	96.71
<b>Average</b>	0.00	40.76	0.79	0.01	10.48	0.00	0.00	20.98	0.21	8.67	11.13	1.65	2.03	96.70
<b>Std Dev.</b>	0.00	0.12	0.03	0.02	0.08	0.00	0.00	0.04	0.01	0.06	0.03	0.02	0.02	0.24
<b>Min</b>	0.00	40.58	0.76	0.00	10.37	0.00	0.00	20.91	0.19	8.59	11.08	1.63	2.02	96.46
<b>Max</b>	0.00	40.95	0.84	0.03	10.59	0.00	0.00	21.02	0.23	8.76	11.17	1.69	2.07	97.24

**Table S13**

Elemental composition of augite as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	50.90	0.76	0.00	3.39	0.00	0.58	5.81	0.11	16.41	19.70	0.33	0.00	97.99
	0.00	51.55	0.64	0.00	2.88	0.00	0.55	5.69	0.13	16.94	19.57	0.34	0.00	98.29
	0.00	51.96	0.55	0.00	2.49	0.00	0.67	5.41	0.11	17.32	19.72	0.32	0.00	98.55
	0.00	51.55	0.73	0.00	2.98	0.03	0.69	5.45	0.12	16.51	20.15	0.33	0.00	98.54
	0.00	51.31	0.77	0.00	3.09	0.06	0.60	5.54	0.12	16.56	20.19	0.37	0.00	98.61
	0.00	51.46	0.76	0.00	3.18	0.00	0.61	5.61	0.13	16.61	20.12	0.35	0.00	98.83
	0.00	51.74	0.75	0.00	3.38	0.03	0.61	5.70	0.11	16.68	20.20	0.33	0.00	99.53
	0.00	52.26	0.75	0.00	3.41	0.00	0.59	5.80	0.13	16.80	20.25	0.37	0.00	100.36
<b>Average</b>	0.00	51.59	0.71	0.00	3.10	0.02	0.61	5.63	0.12	16.73	19.99	0.34	0.00	98.84
<b>Std Dev.</b>	0.00	0.41	0.08	0.00	0.32	0.02	0.05	0.15	0.01	0.29	0.27	0.02	0.00	0.76
<b>Min</b>	0.00	50.90	0.55	0.00	2.49	0.00	0.55	5.41	0.11	16.41	19.57	0.32	0.00	97.99
<b>Max</b>	0.00	52.26	0.77	0.00	3.41	0.06	0.69	5.81	0.13	17.32	20.25	0.37	0.00	100.36

**Table S14**

Elemental composition of epidote as obtained via electron microprobe analysis. In all cases, results less than the estimated limit of detection are reported as zero.

	Nb <sub>2</sub> O <sub>5</sub>	SiO <sub>2</sub>	TiO <sub>2</sub>	ZnO	Al <sub>2</sub> O <sub>3</sub>	V <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	Total
	0.00	37.19	0.22	0.00	22.39	0.05	0.00	14.18	0.14	0.03	23.03	0.00	0.00	97.23
	0.00	37.03	0.23	0.00	22.62	0.04	0.00	14.13	0.14	0.03	23.16	0.00	0.00	97.38
	0.00	37.31	0.17	0.00	22.66	0.05	0.00	14.24	0.14	0.03	23.01	0.00	0.00	97.61
	0.00	37.02	0.23	0.00	22.60	0.04	0.00	14.15	0.13	0.04	23.02	0.00	0.00	97.23
	0.00	37.20	0.20	0.00	22.63	0.05	0.00	14.18	0.12	0.03	23.03	0.00	0.00	97.44
	0.00	37.13	0.20	0.00	22.49	0.04	0.00	14.18	0.11	0.05	23.07	0.00	0.00	97.27
	0.00	37.20	0.20	0.00	22.53	0.04	0.00	14.23	0.12	0.03	23.13	0.00	0.00	97.48
	0.00	37.19	0.22	0.00	22.38	0.05	0.00	14.34	0.12	0.02	23.13	0.00	0.00	97.45
<b>Average</b>	0.00	37.16	0.21	0.00	22.54	0.05	0.00	14.20	0.13	0.03	23.07	0.00	0.00	97.39
<b>Std Dev.</b>	0.00	0.10	0.02	0.00	0.11	0.01	0.00	0.07	0.01	0.01	0.06	0.00	0.00	0.14
<b>Min</b>	0.00	37.02	0.17	0.00	22.38	0.04	0.00	14.13	0.11	0.02	23.01	0.00	0.00	97.23
<b>Max</b>	0.00	37.31	0.23	0.00	22.66	0.05	0.00	14.34	0.14	0.05	23.16	0.00	0.00	97.61

**Table S15.** Crystallographic data for Ti-containing sample refinement.

sample	Commercial anatase	Anatase 1	Anatase 2	Rutile
space group	$I4_1/amd$ (No. 136)	$I4_1/amd$ (No. 136)	$I4_1/amd$ (No. 136)	$P4_2/mmm$ (No. 136)
$a$ (Å)	3.78645(5)	3.78514(4)	3.78489(3)	4.59572(6)
$b$ (Å)	3.78645(5)	3.78514(4)	3.78489(3)	4.59572(6)
$c$ (Å)	9.5179(1)	9.5120(1)	9.5119(1)	2.95966(4)
$\alpha$ °	90	90	90	90
$\beta$ °	90	90	90	90
$\gamma$ °	90	90	90	90
$T$ (K)	296	296	296	296
radiation	Co $K\alpha$ ,	Co $K\alpha$ ,	Co $K\alpha$ ,	Co $K\alpha$ ,
	K $\alpha$ 1 $\lambda$ =1.78900 Å			
	K $\alpha$ 2 $\lambda$ =1.79283 Å			
$2\theta$ limits	20.00–90.00°	20.00–90.00°	20.00–90.00°	20.00–90.00°
refinement method	Pawley	Pawley	Pawley	Pawley
no. of data collected	3500 data points	3500 data points	3500 data points	3500 data points
no. of Bragg reflections	14	14	14	13
no. of variables	27	27	27	27
residuals	$R_{wp} = 0.0472$	$R_{wp} = 0.0492$	$R_{wp} = 0.0657$	$R_{wp} = 0.0482$

$$R_{wp} = \left[ \frac{\sum [w(y_o - y_c)]}{\sum w y_o^2} \right]^{1/2}$$

sample	Titanite 1	Titanite 2	Phlogopite	Augite	Hastingsite	Epidote	ilmeneite
space group	C2/c (No. 15)	C2/c (No. 15)	C2/m (No. 12)	C2/c (No. 15)	C2/m (No. 12)	P2 <sub>1</sub> /m (No. 11)	R $\bar{3}$ (No. 148)
a (Å)	6.5529(1)	6.5735(1)	5.1635(2)	9.7361(2)	9.9076(1)	8.8986(1)	5.08769(7)
b (Å)	8.7022(2)	8.7278(1)	9.0833(6)	8.8989(1)	18.1624(4)	5.63531(8)	5.08769(7)
c (Å)	7.0569(1)	7.0851(1)	20.086(2)	5.2686(1)	5.3283(3)	10.1623(1)	4.0820(2)
$\alpha$ °	90	90	90	90	90	90	90
$\beta$ °	113.849(1)	113.962(1)	95.685(3)	106.327(1)	104.992(3)	115.4145(8)	90
$\gamma$ °	90	90	90	90	90	90	120
T (K)	296	296	296	296	296	296	296
radiation	Co K $\alpha$ ,						
	K $\alpha$ 1 $\lambda$ =1.78900 Å						
2 $\theta$ limits	K $\alpha$ 2 $\lambda$ =1.79283 Å 20.00-90.00°						
refinement method	Pawley						
no. of data collected	3500 data points						
no. of Bragg reflections	189	189	257	14	258	281	36
no. of variables	205	205	273	27	330	333	50
residuals	$R_{wp} = 0.0731$	$R_{wp} = 0.0482$	$R_{wp} = 0.0923$	$R_{wp} = 0.0526$	$R_{wp} = 0.0347$	$R_{wp} = 0.0358$	$R_{wp} = 0.0201$

$$R_{wp} = \left[ \frac{\sum [w(y_o - y_c)]}{\sum w y_o^2} \right]^{1/2}$$

**Table S16.** Crystallographic data for anatase 2 and titanite 1 single crystals

Formula	TiO <sub>2</sub>	CaTiSiO <sub>5</sub>
Formula mass (amu)	79.90	196.07
Space group	<i>I4<sub>1</sub>/amd</i> (No. 136)	<i>C2/c</i> (No. 15)
<i>a</i> (Å)	3.806(4)	6.572(3)
<i>b</i> (Å)	3.806(4)	8.731(3)
<i>c</i> (Å)	9.576(9)	7.084(4)
$\alpha$ , °	90	90
$\beta$ , °	90	113.922(4)
$\gamma$ , °	90	90
<i>V</i> (Å <sup>3</sup> )	138.7(2)	371.6(3)
<i>Z</i>	4	4
$\rho_{\text{calcd}}$ (g cm <sup>-3</sup> )	3.825	3.505
<i>T</i> (K)	296(2)	296(2)
Crystal dimensions (mm)	0.01 × 0.08 × 0.20	0.05 × 0.12 × 0.14
Radiation	Graphite monochromated Mo <i>K</i> $\alpha$ , $\lambda = 0.71073$ Å	
$\mu$ (Mo <i>K</i> $\alpha$ ) (mm <sup>-1</sup> )	5.55	3.900
Transmission factors	0.3958–0.9313	0.6191–0.8441
$2\theta$ limits	11.54 – 95.74 °	8.24 – 66.44 °
Data collected	$-5 \leq h \leq 5, -5 \leq k \leq 5, -$ $14 \leq l \leq 14$	$-9 \leq h \leq 10, -13 \leq k \leq 13, -$ $10 \leq l \leq 10$
No. of data collected	917	2164
No. of unique data, including $F_o^2 < 0$	84 ( $R_{\text{int}} = 0.0370$ )	706 ( $R_{\text{int}} = 0.0285$ )
No. of unique data, with $F_o^2 > 2\sigma(F_o^2)$	72	560
No. of variables	8	41
$R(F)$ for $F_o^2 > 2\sigma(F_o^2)$ <sup>a</sup>	0.0127	0.0373
$R_w(F_o^2)$ <sup>b</sup>	0.0144	0.0481
Goodness of fit	1.231	1.084
$(\Delta\rho)_{\text{max}}, (\Delta\rho)_{\text{min}}$ (e Å <sup>-3</sup> )	0.362, -0.448	0.644, -1.262

<sup>a</sup>  $R(F) = \sum ||F_o| - |F_c|| / \sum |F_o|$ . <sup>b</sup>  $R_w(F_o^2) = [\sum [w(F_o^2 - F_c^2)^2] / \sum wF_o^4]^{1/2}$ ;  $w^{-1} = [\sigma^2(F_o^2) + (Ap)^2 + Bp]$ , where  $p = [\max(F_o^2, 0) + 2F_c^2] / 3$ .

**Table S17.** Atomic coordinates and displacement parameters for anatase 2.

Position	Wyckoff position	<i>x</i>	<i>y</i>	<i>z</i>	$U_{11}$ (Å <sup>2</sup> )	$U_{22}$ (Å <sup>2</sup> )	$U_{33}$ (Å <sup>2</sup> )
Ti	4 <i>b</i>	0	1/4	3/8	0.0081(3)	0.0081(3)	0.0059(3)
O	8 <i>e</i>	0	1/4	0.1668(2)	0.0069(6)	0.0134(6)	0.0064(6)

**Table S18.** Selected interatomic distances (Å) for anatase 2.

Ti - O (×4)	1.945(2)
Ti - O (×2)	1.993(2)
Ti - Ti (×4)	3.058(2)

**Table S19.** Atomic coordinates and displacement parameters for titanite 1.

Position	Wyckoff position	<i>x</i>	<i>y</i>	<i>z</i>	<i>U</i> <sub>11</sub> (Å <sup>2</sup> )	<i>U</i> <sub>22</sub> (Å <sup>2</sup> )	<i>U</i> <sub>33</sub> (Å <sup>2</sup> )
Ti	4 <i>a</i>	0	0	0	0.0102(3)	0.0118(3)	0.0225(4)
Si	4 <i>e</i>	0	0.68280(9)	1/4	0.0090(4)	0.0091(4)	0.0111(4)
Ca	4 <i>e</i>	0	0.33202(7)	1/4	0.0120(3)	0.0103(3)	0.0361(5)
O1	4 <i>e</i>	0	0.0721(3)	1/4	0.0174(10)	0.0122(9)	0.0123(10)
O2	8 <i>f</i>	0.3976(2)	0.28987(16)	0.3819(2)	0.0144(7)	0.0121(7)	0.0183(8)
O3	8 <i>f</i>	0.3141(2)	0.06639(19)	0.0897(2)	0.0141(7)	0.0147(8)	0.0177(8)

**Table S20.** Selected interatomic distances (Å) for titanite 1.

---

Ti - O1 (×2)	1.879(1)
Ti - O3 (×2)	1.984(2)
Ti - O2 (×2)	2.016(2)
Si - O3 (×2)	1.640(2)
Si - O2 (×2)	1.646(2)
Ca - O1	2.269(3)
Ca - O3 (×2)	2.417(2)
Ca - O2 (×2)	2.420(2)

---