

Organic complexity in protostellar disk candidates: Supporting information

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Moment zero map parameters

Table 1 lists the rms values and velocity ranges used to make the moment zero maps shown in Figures 1 and 2 in the main text.

Table 1: Velocity ranges and rms values used for the integrated intensity maps (Figures 1 and 2) in the main text. For clarity, only upper state quantum numbers are used to identify each line; refer to Table 3 in the main text for full identifiers.

	Ser-emb 1		Ser-emb 7		Ser-emb 8		Ser-emb 15		Ser-emb 17	
	Vel. (km/s)	rms (mJy)								
C ¹⁸ O 2	6.7–10.5	4.1	6.0–12.0	5.4	6.0–10.7	5.6	8.2–12.5	4.4	4.7–10.5	5.7
CH ₃ OH 5 _{1,4}	2.6–11.6	6.8	6.2–11.6	4.1	2.6–14.0	6.5	8.6–12.2	3.2	2.0–13.4	7.5
CH ₃ OH 10 _{2,8}	7.0–10.2	2.9	4.5–12.7	5.4	5.1–10.8	3.5
CH ₃ OCH ₃ 13 _{0,13}	6.4–12.1	3.7	3.2–12.1	5.5	4.5–12.1	4.3
CH ₃ OCHO 19 _{4,15}	5.7–9.5	3.4	5.7–12.6	5.0	4.4–10.1	4.1
NH ₂ CHO 12 _{1,12}	6.8–9.8	3.2	6.2–12.8	5.2	5.0–12.8	5.3
CH ₂ CO 12 _{1,11}	5.6–12.1	4.7	4.4–12.1	5.4	4.4–12.1	5.1

Spectral line fits

Figures 1–4 show Gaussian fits to the observed lines of each COM, analogous to Figure 3 in the main text.

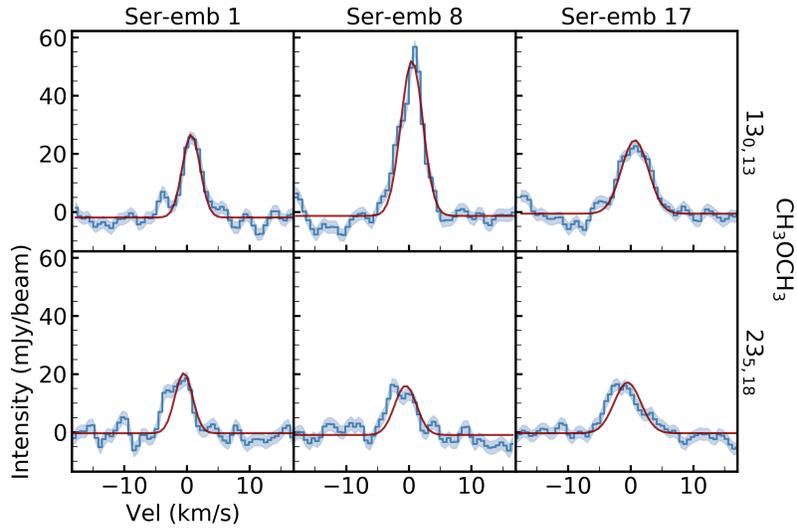


Figure 1: CH_3OCH_3 spectral lines. Blue lines show the spectra extracted from the continuum peak pixel, and shaded regions represent the rms. Red lines show Gaussian fits to the data.

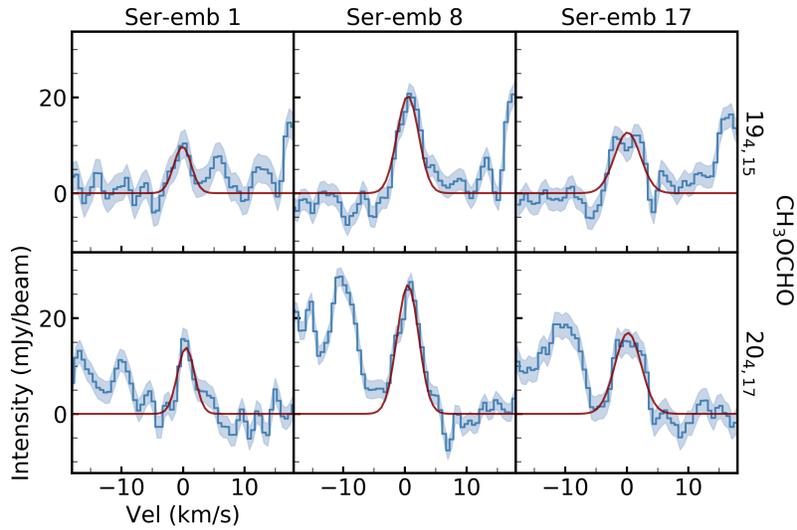


Figure 2: CH_3OCHO spectral lines. Blue lines show the spectra extracted from the continuum peak pixel, and shaded regions represent the rms. Red lines show Gaussian fits to the data.

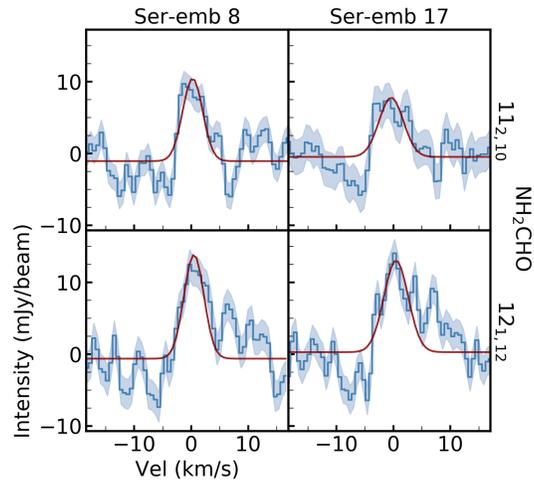


Figure 3: NH₂CHO spectral lines. Blue lines show the spectra extracted from the continuum peak pixel, and shaded regions represent the rms. Red lines show Gaussian fits to the data.

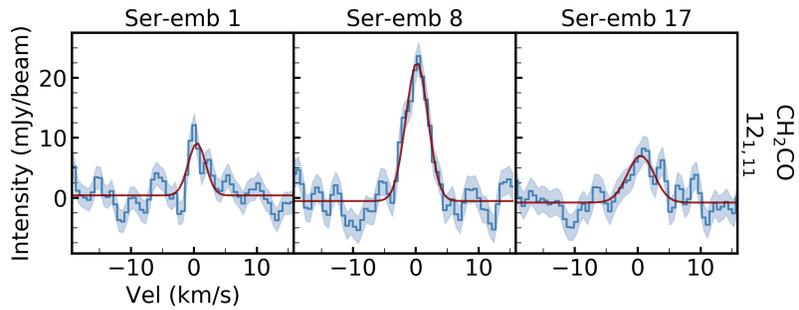


Figure 4: CH₂CO spectra lines. Blue lines show the spectra extracted from the continuum peak pixel, and shaded regions represent the rms. Red lines show Gaussian fits to the data.

Full spectra

Figures 5–8 show the full spectra extracted from the continuum peak pixel in Ser-emb 1, 7, 8, and 15, analogous to Figure 5 in the main text. For Ser-emb 1 and 8 (Figures 5 and 7) colored lines show the synthetic spectra of COMs detected in each source.

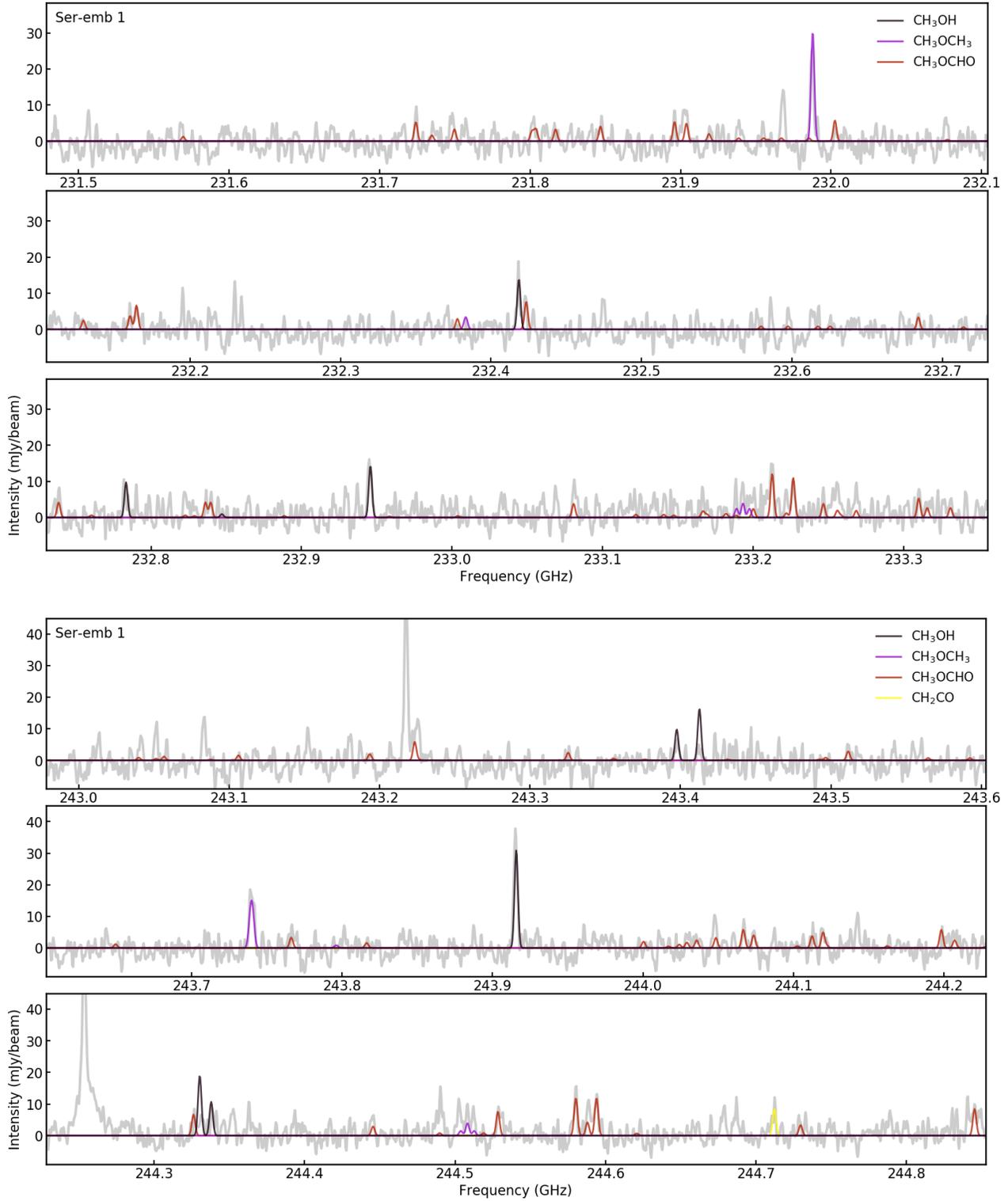


Figure 5: Full spectrum extracted from the continuum peak pixel in Ser-emb 1 (grey line), along with synthetic spectra of the detected COMs (colored lines). Spectra are calculated assuming the CH₃OH rotational temperature.

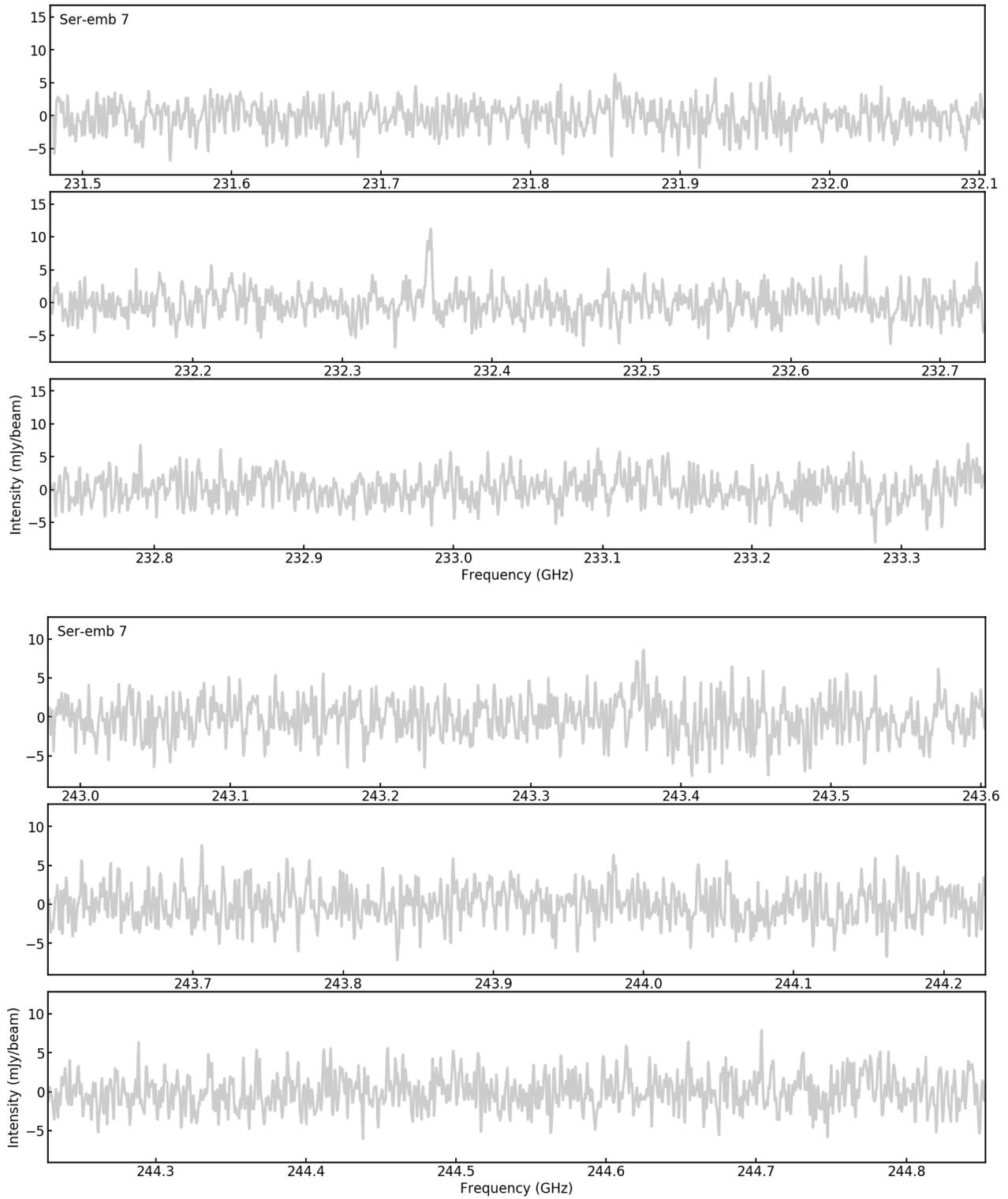


Figure 6: Full spectrum extracted from the continuum peak pixel in Ser-emb 7 (grey line).

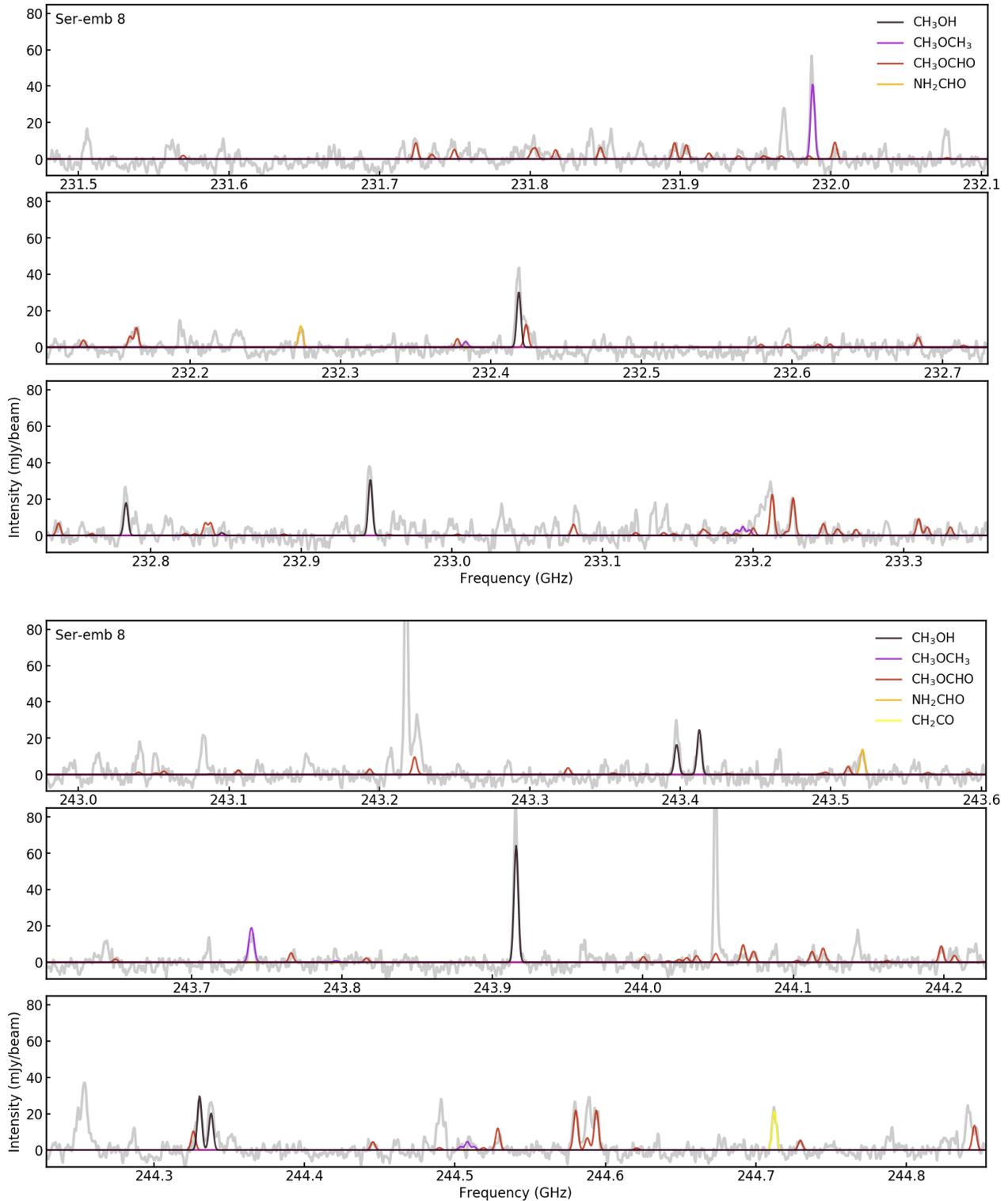


Figure 7: Full spectrum extracted from the continuum peak pixel in Ser-emb 8 (grey line), along with synthetic spectra of the detected COMs (colored lines). Spectra are calculated assuming the CH₃OH rotational temperature.

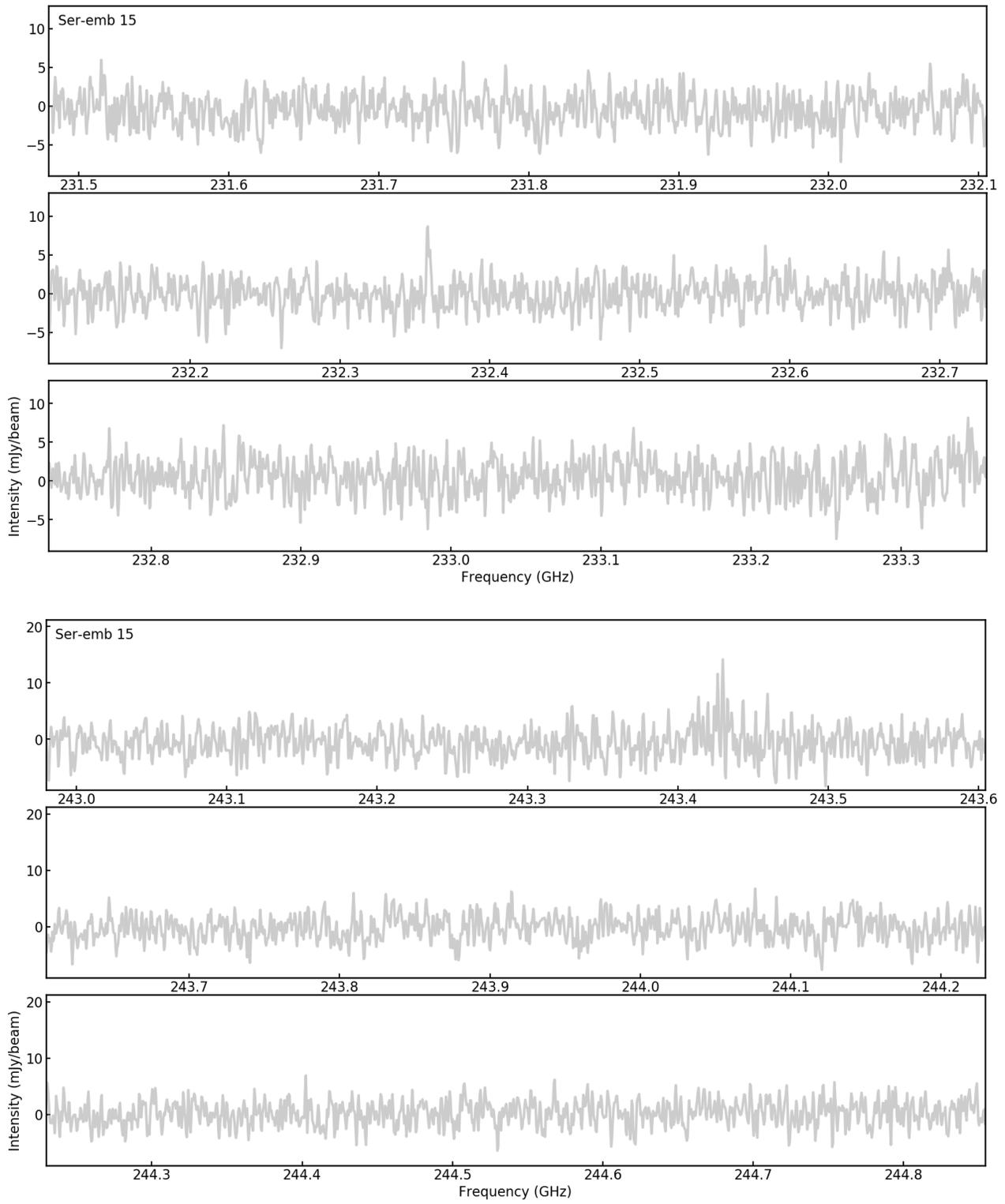


Figure 8: Full spectrum extracted from the continuum peak pixel in Ser-emb 15 (grey line).

Population diagram diagram fitting

For the MCMC population diagram fits to CH₃OH, we use a flat prior $10^5 < N_T < 10^{20}$ cm⁻² and $100 < T_r < 400$ K. 200 walkers are propagated for 1000 steps, and the samples are well converged. Walker chains and corner plots for each source are shown in Figures 9–11 for the maximum beam dilution case.

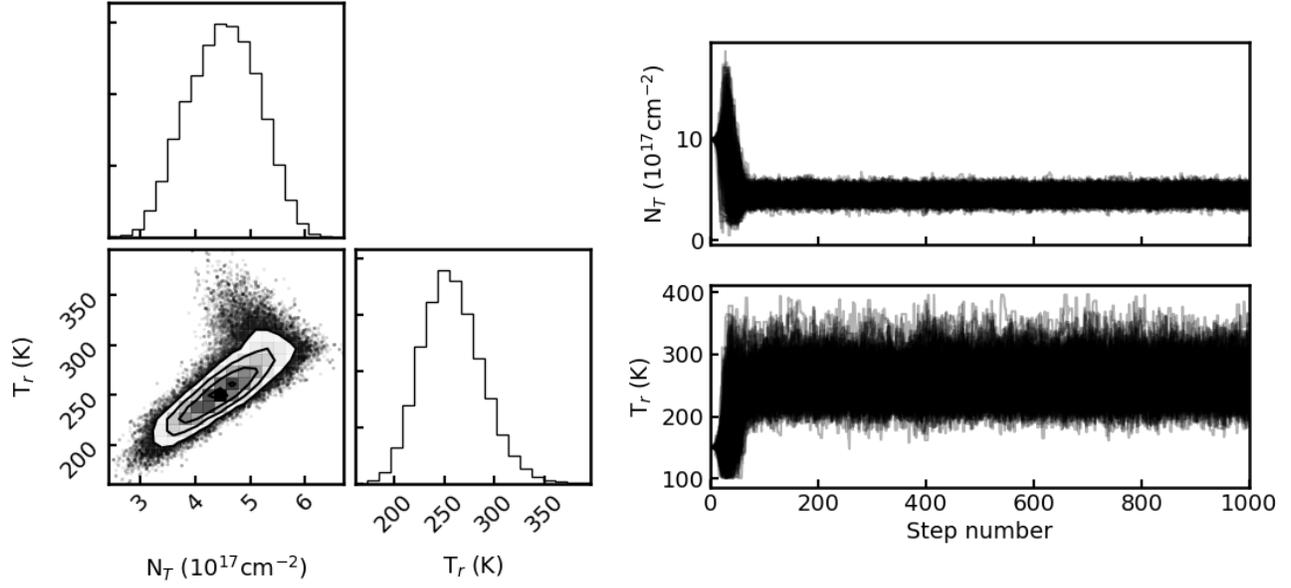


Figure 9: Ser-emb 1 rotational diagram MCMC fit results. The corner plot is shown on the left and the walker chain on the right.

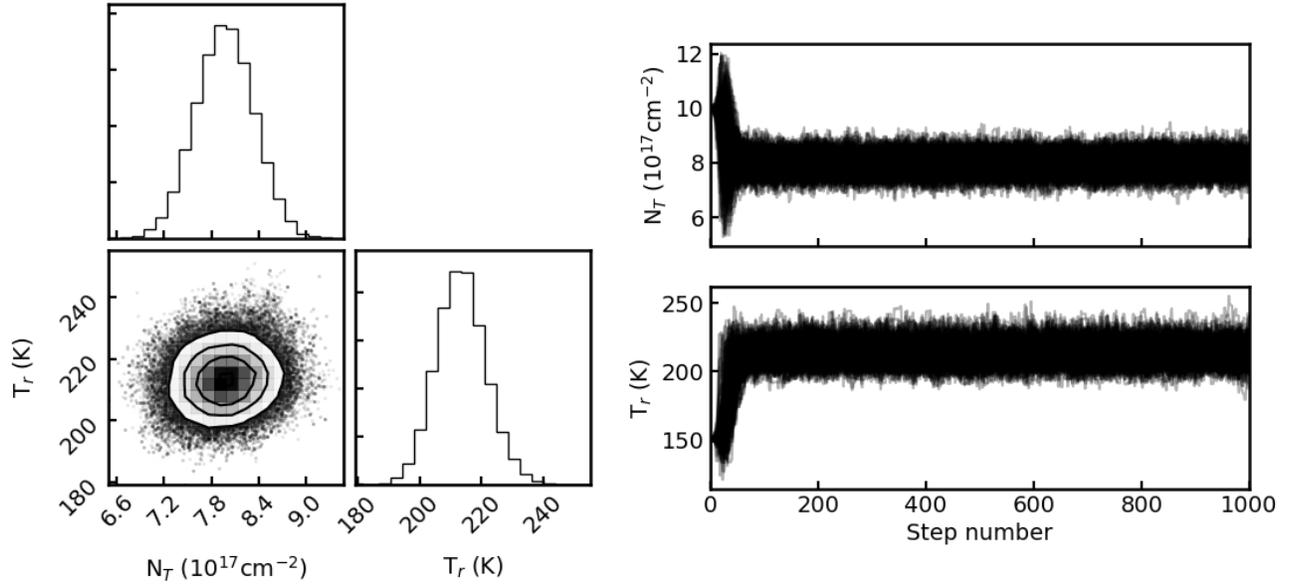


Figure 10: Ser-emb 8 rotational diagram MCMC fit results. The corner plot is shown on the left and the walker chain on the right.

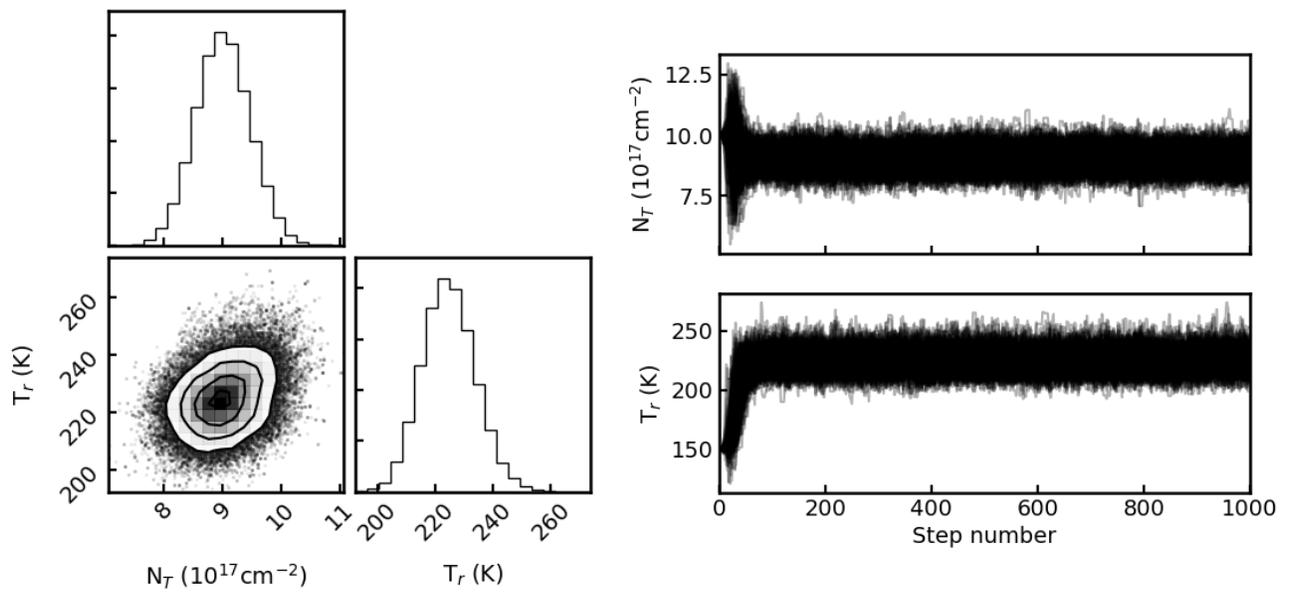


Figure 11: Ser-emb 17 rotational diagram MCMC fit results. The corner plot is shown on the left and the walker chain on the right.